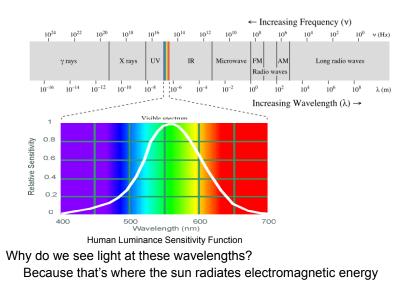
Color and Radiometry

Digital Image Synthesis Yung-Yu Chuang

with slides by Svetlana Lazebnik, Pat Hanrahan and Matt Pharr

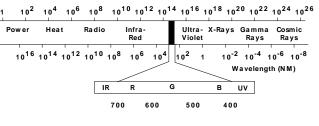
Electromagnetic spectrum





Radiometry

- Radiometry: study of the propagation of electromagnetic radiation in an environment
- Four key quantities: flux, intensity, irradiance and radiance
- These radiometric quantities are described by their spectral power distribution (SPD)
- Human visible light ranges from 370nm to 730nm



Basic radiometry



- pbrt is based on radiative transfer: study of the transfer of radiant energy based on radiometric principles and operates at the geometric optics level (light interacts with objects much larger than the light's wavelength)
- It is based on the particle model. Hence, diffraction and interference can't be easily accounted for.



Basic assumptions about light behavior

- Linearity: the combined effect of two inputs is equal to the sum of effects
- Energy conservation: scattering event can't produce more energy than they started with
- Steady state: light is assumed to have reached equilibrium, so its radiance distribution isn't changing over time.
- No polarization: we only care the frequency of light but not other properties (such as phases)
- No fluorescence or phosphorescence: behavior of light at a wavelength or time doesn't affect the behavior of light at other wavelengths or time

Fluorescent materials

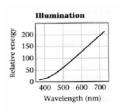


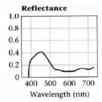


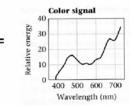
Interaction of light and surfaces



*







Reflected color is the result of interaction of light source spectrum with

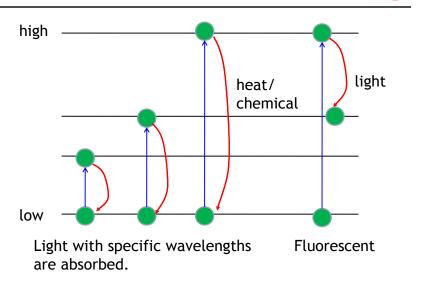
surface reflectance

- All definitions and units are now "per unit wavelength" - All terms are now "spectral"

Spectral radiometry

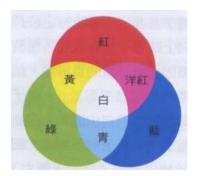
Color

Why reflecting different colors



Primary colors





Primary colors for addition (light sources)

Primary colors for subtraction (reflection)

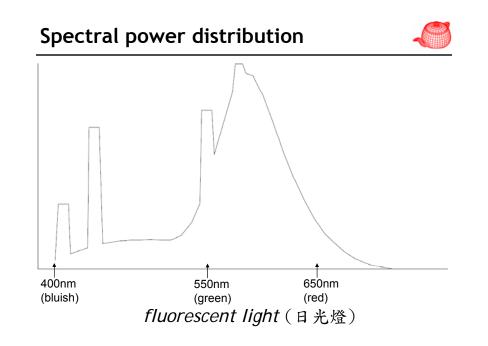
黃

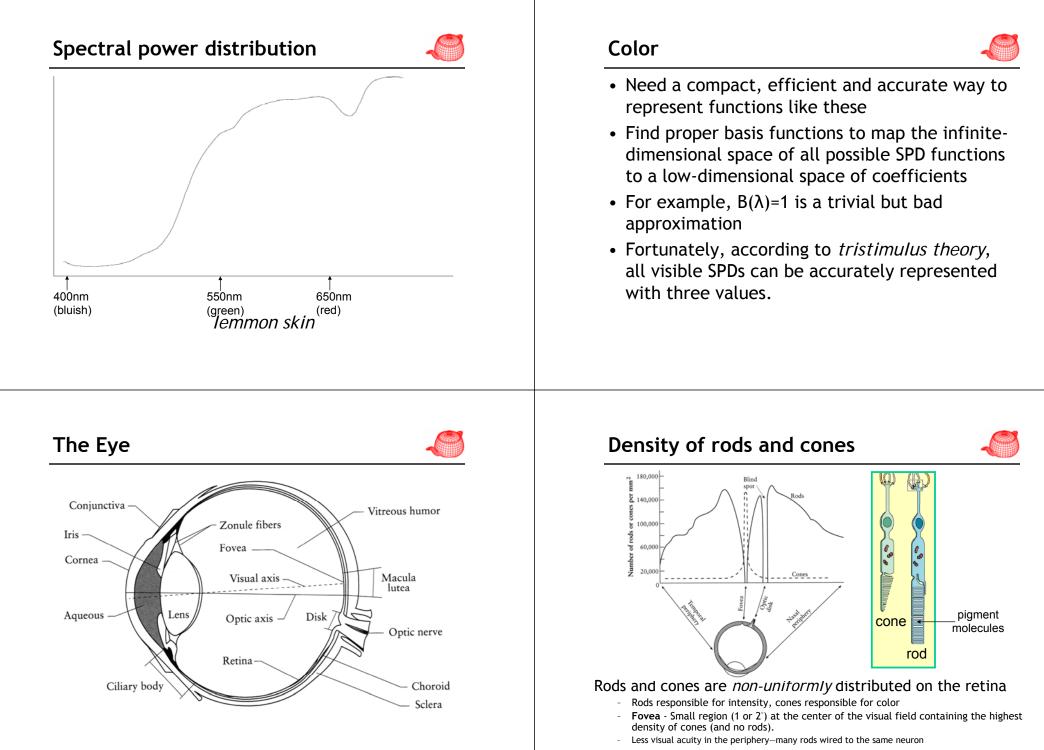
洋紅

Heat generates light



- Vibration of atoms or electrons due to heat generates electromagnetic radiation as well. If its wavelength is within visible light (>1000K), it generates color as well.
- Color only depends on temperature, but not property of the object.
- Human body radiates IR light under room temperature.
- 2400-2900K: color temperature of incandescent light bulb

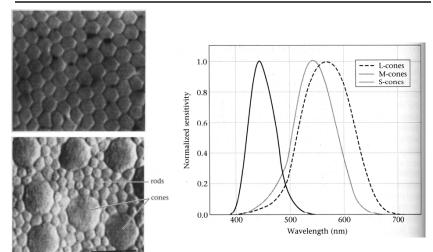


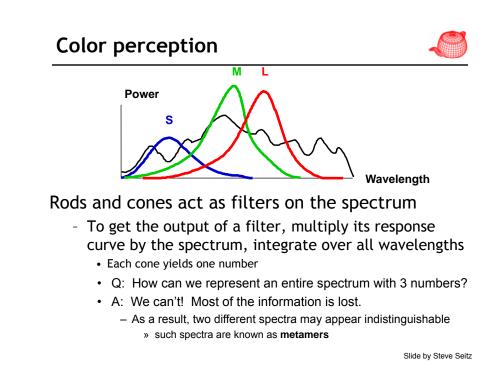


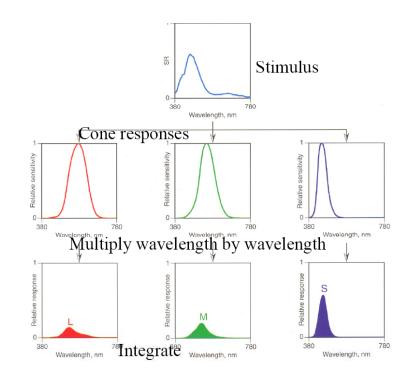
Slide by Steve Seitz

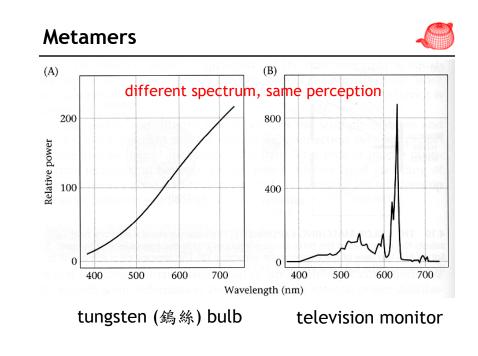
Human Photoreceptors

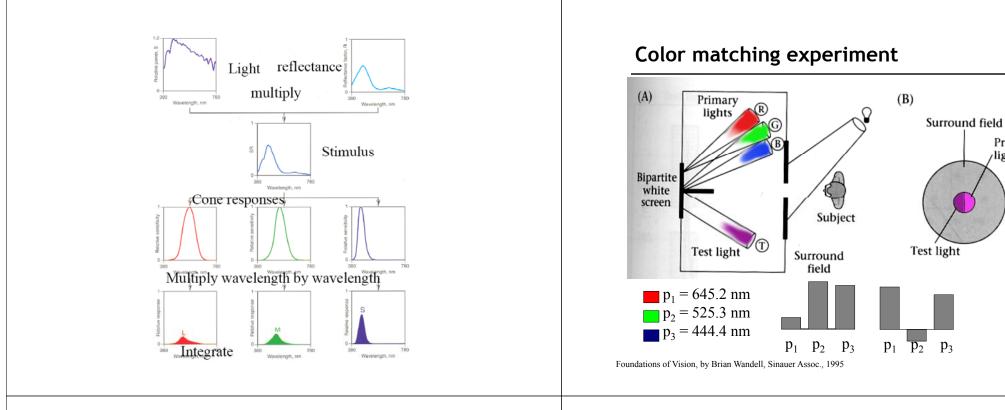




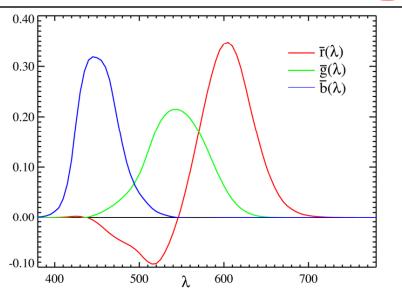








Color matching experiment



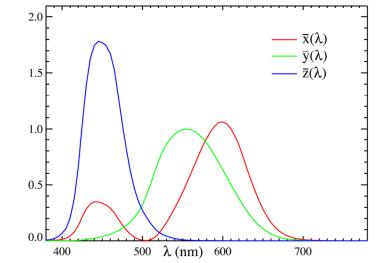
Color matching experiment



Primary

/lights

• To avoid negative parameters



Spectrum



- In core/spectrum.*
- Two representations: RGBSpectrum (default) and SampledSpectrum
- The selection is done at compile time with a typedef in core/pbrt.h
 typedef RGBSpectrum Spectrum;
- Both stores a fixed number of samples at a fixed set of wavelengths.

CoefficientSpectrum



template <int nSamples>
class CoefficientSpectrum {
 +=, +, -, /, *, *= (CoefficientSpectrum)
 ==, != (CoefficientSpectrum)
 IsBlack, Clamp
 *, *=, /, /= (float)
protected:
 float c[nSamples];
}

Sqrt, Pow, Exp

SampledSpectrum



• Represents a SPD with uniformly spaced samples between a starting and an ending wavelength (400 to 700 nm for HVS). The number of samples, 30, is generally more than enough.

static const int sampledLambdaStart = 400; static const int sampledLambdaEnd = 700; static const int nSpectralSamples = 30;

SampledSpectrum



class SampledSpectrum : public CoefficientSpectrum<nSpectralSamples> {

}

It is possible to convert SPD with irregular spaced samples and more or fewer samples into a SampledSpectrum. For example, sampled BRDF.

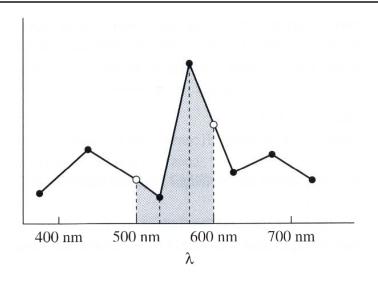
SampledSpectrum



```
static SampledSpectrum FromSampled(
        float *lambda, float *v, int n) {
  <Sort samples if unordered>
  SampledSpectrum r;
 for (int i = 0; i<nSpectralSamples; ++i) {</pre>
   lambda0=Lerp(i/float(nSpectralSamples),
      sampledLambdaStart, sampledLambdaEnd);
   lambda1=Lerp((i+1)/float(nSpectralSamples),
      sampledLambdaStart, sampledLambdaEnd);
   r.c[i]=AverageSpectrumSamples(lambda,
           v, n, lambda0, lambda1);
 return r;
```

AverageSpectrumSamples





Human visual system



- Tristimulus theory: all visible SPDs S can be accurately represented for human observers with three values, x_{λ} , y_{λ} and z_{λ} .
- The basis are the *spectral matching curves*, $X(\lambda)$, $Y(\lambda)$ and $Z(\lambda)$ determined by CIE (國際照明委員 會).

$$x_{\lambda} = \int_{\lambda} S(\lambda) X(\lambda) d\lambda$$
$$y_{\lambda} = \int_{\lambda} S(\lambda) Y(\lambda) d\lambda$$
$$z_{\lambda} = \int_{\lambda} S(\lambda) Z(\lambda) d\lambda$$

XYZ basis pbrt has discrete versions (sampled every 1nm) 2 of these bases in core/color.cpp X --- Y 1.5 Z 0.5 450 500 550 600 650 700 400 360

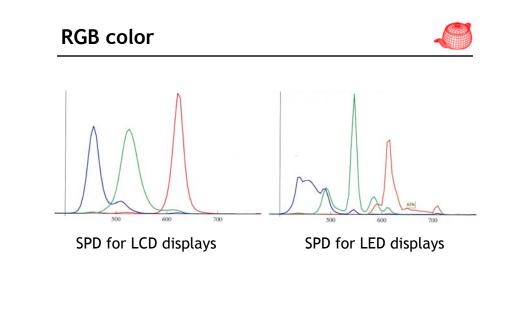
XYZ color

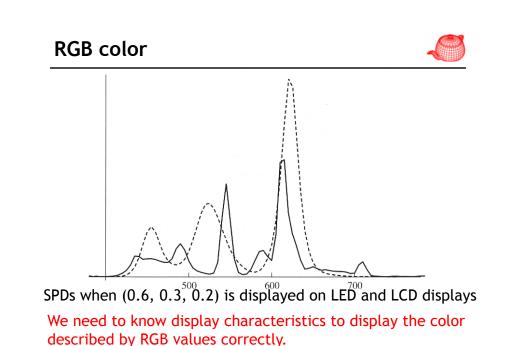


- Good for representing visible SPD to human observer, but not good for spectral computation.
- A product of two SPD's XYZ values is likely different from the XYZ values of the SPD which is the product of the two original SPDs.
- It is frequent to convert our samples into XYZ
- In Init(), we initialize the following static SampledSpectrum X, Y, Z;
- static float yint; x.c[i] stores the sum of X function
 yint stores the within the ith wavelength interval
 - sum of Y.c[i] using AverageSpectrumSamples

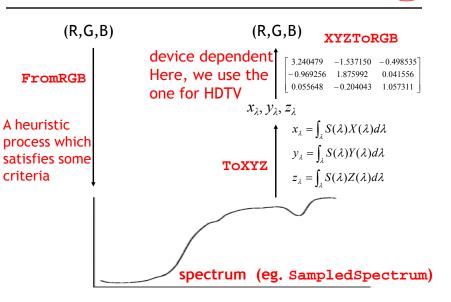
XYZ color

```
void ToXYZ(float xyz[3]) const {
    xyz[0] = xyz[1] = xyz[2] = 0.;
    for (int i = 0; i < nSpectralSamples; ++i)
    {
        xyz[0] += X.c[i] * c[i];
        xyz[1] += Y.c[i] * c[i];
        xyz[2] += Z.c[i] * c[i];
    }
    xyz[0] /= yint;
    xyz[0] /= yint;
    xyz[0] /= yint;
}</pre>
```





Conversions



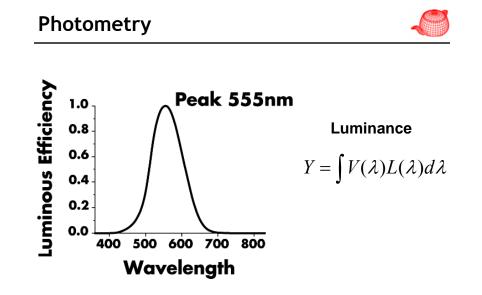
Radiometry





 Note that RGB representation is ill-defined. Same RGB values display different SPDs on different displays. To use RGB to display a specific SPD, we need to know display characteristics first. But, it is convenient, computation and storage efficient.

class RGBSpectrum : public CoefficientSpectrum<3> { using CoefficientSpectrum<3>::c;



Basic quantities



non-directional

Flux: power, (W)

Irradiance: flux density per area, (W/m²)

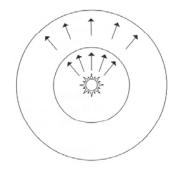
directional

Intensity: flux density per solid angle Radiance: flux density per solid angle per area

Flux (Φ)

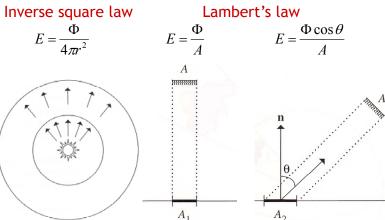


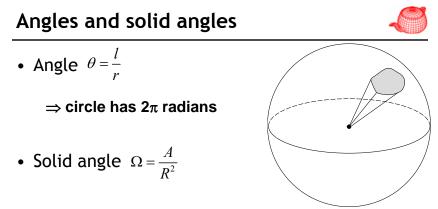
- Radiant flux, power
- Total amount of energy passing through a surface per unit of time (J/s,W)





• Area density of flux (W/m²) $E = \frac{d\Phi}{dA}$





The solid angle subtended by a surface is defined as the surface area of a unit sphere covered by the surface's projection onto the sphere.

 \Rightarrow sphere has 4π steradians

Intensity (I)

 $I(\omega) \equiv \frac{d\Phi}{d\Phi}$

dω



- Flux density per solid angle $I = \frac{d\Phi}{d}$
- Intensity describes the directional distribution of light



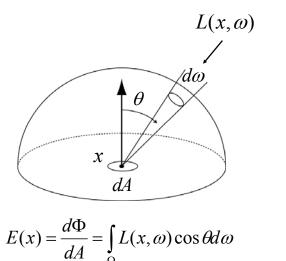
• Flux density per unit area per solid angle

dA

 $L = \frac{d\Phi}{d\omega dA^{\perp}}$

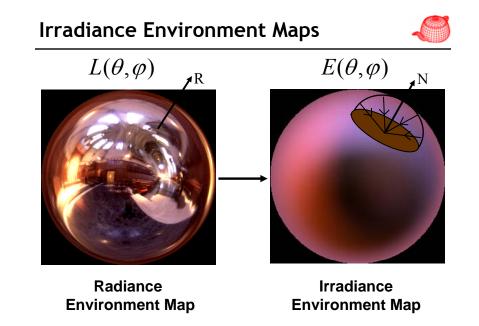
- Most frequently used, remains constant along ray.
- All other quantities can be derived from radiance

Calculate irradiance from radiance





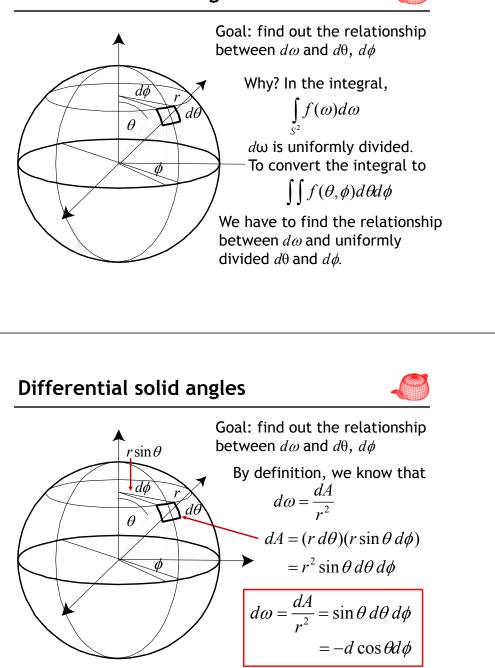
Light meter





Differential solid angles





Differential solid angles

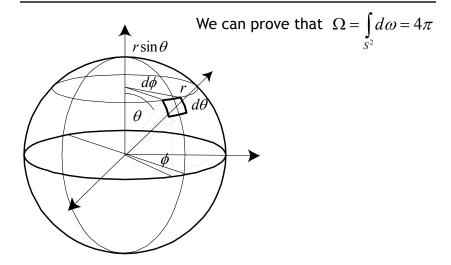


• Can we find the surface area of a unit sphere by $\int_0^{2\pi} \int_0^{\pi} d\theta d\phi$?

 $\int_0^{2\pi} \int_0^{\pi} d\theta d\phi =$

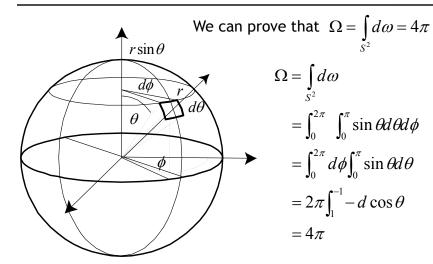
Differential solid angles

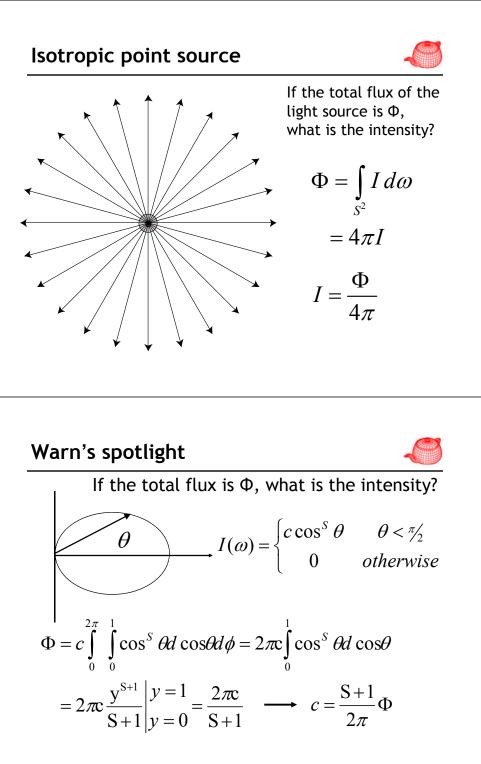




Differential solid angles







Warn's spotlight



```
If the total flux is \Phi, what is the intensity?

\theta
I(\omega) \propto \cos^{S} \theta
```