

Color and Radiometry

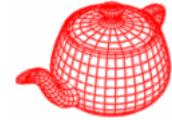
Digital Image Synthesis

Yung-Yu Chuang

11/01/2005

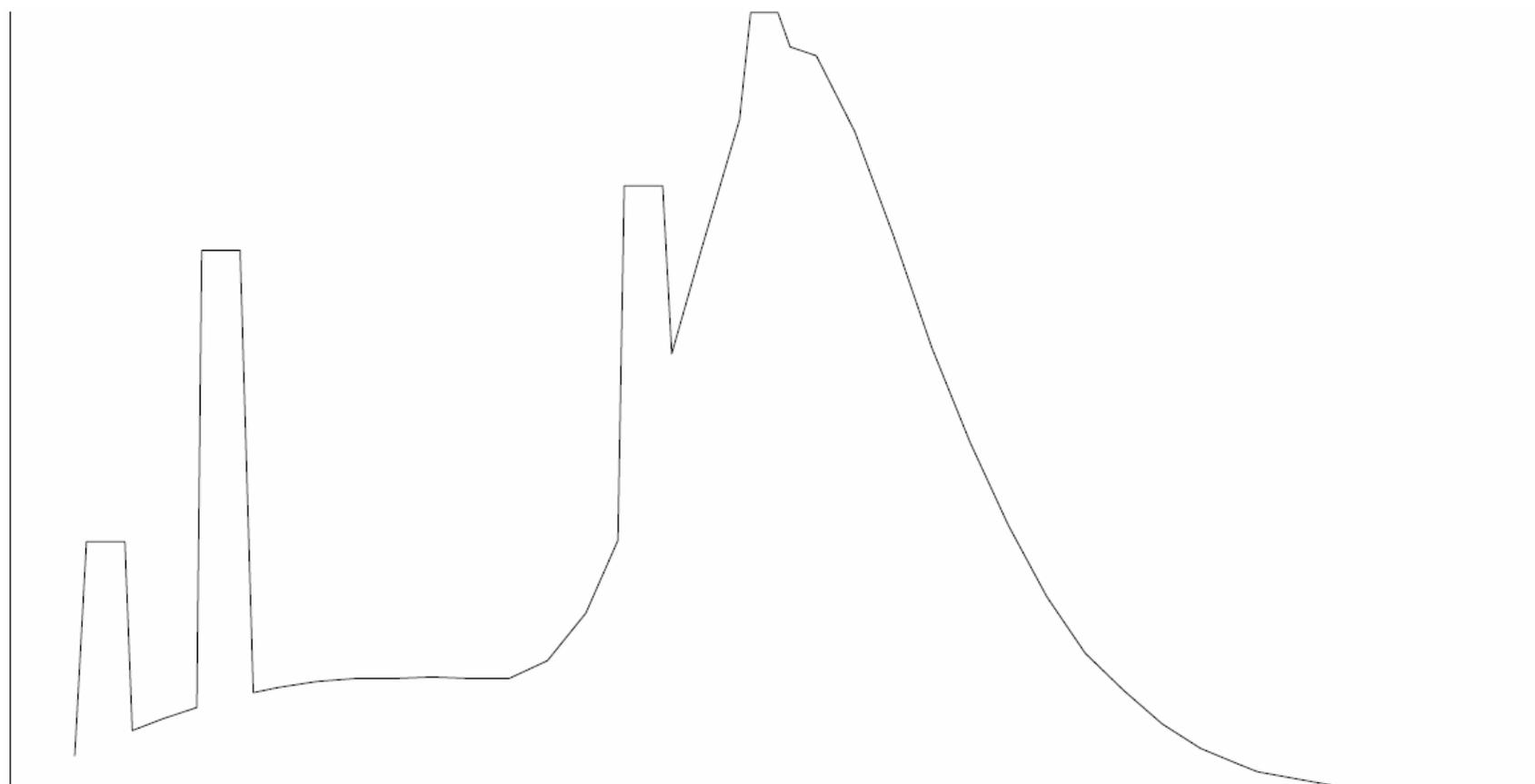
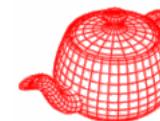
with slides by Pat Hanrahan and Matt Pharr

Radiometry



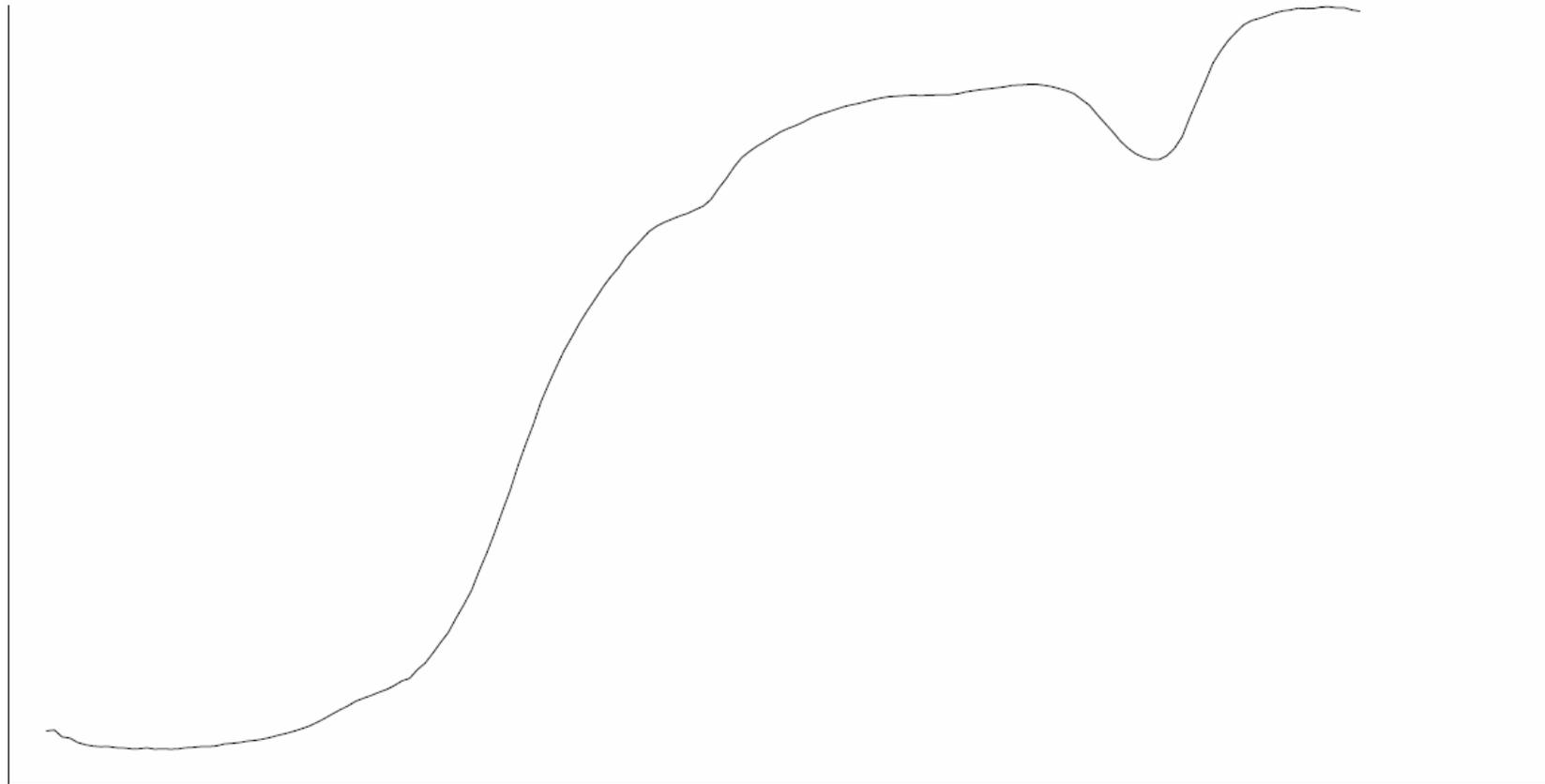
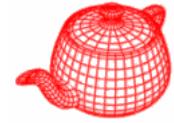
- Radiometry: study of the propagation of electromagnetic radiation in an environment
- Four key quantities: flux, intensity, irradiance and radiance
- These radiometric quantities are described by their spectral power distribution (SPD)

Spectral representation



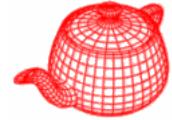
fluorescent light

SPD



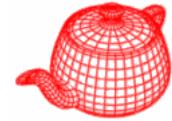
lemmon skin

Color



-
- Need a compact, efficient and accurate way to represent functions like these
 - Find proper basis functions to map the infinite-dimensional space of all possible SPD functions to a low-dimensional space of coefficients
 - For example, $B(\lambda)=1$, bad approximation

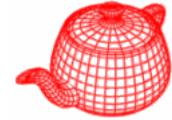
Spectrum



- In `core/color.*`
- Not a plug-in, to use inline for performance
- `Spectrum` stores a fixed number of samples at a fixed set of wavelengths. Better for smooth functions.

```
#define COLOR_SAMPLE 3
class COREDLL Spectrum {
public:
    <arithmetic operations>
private:
    float c[COLOR_SAMPLES];
    ...
}
```

Human visual system

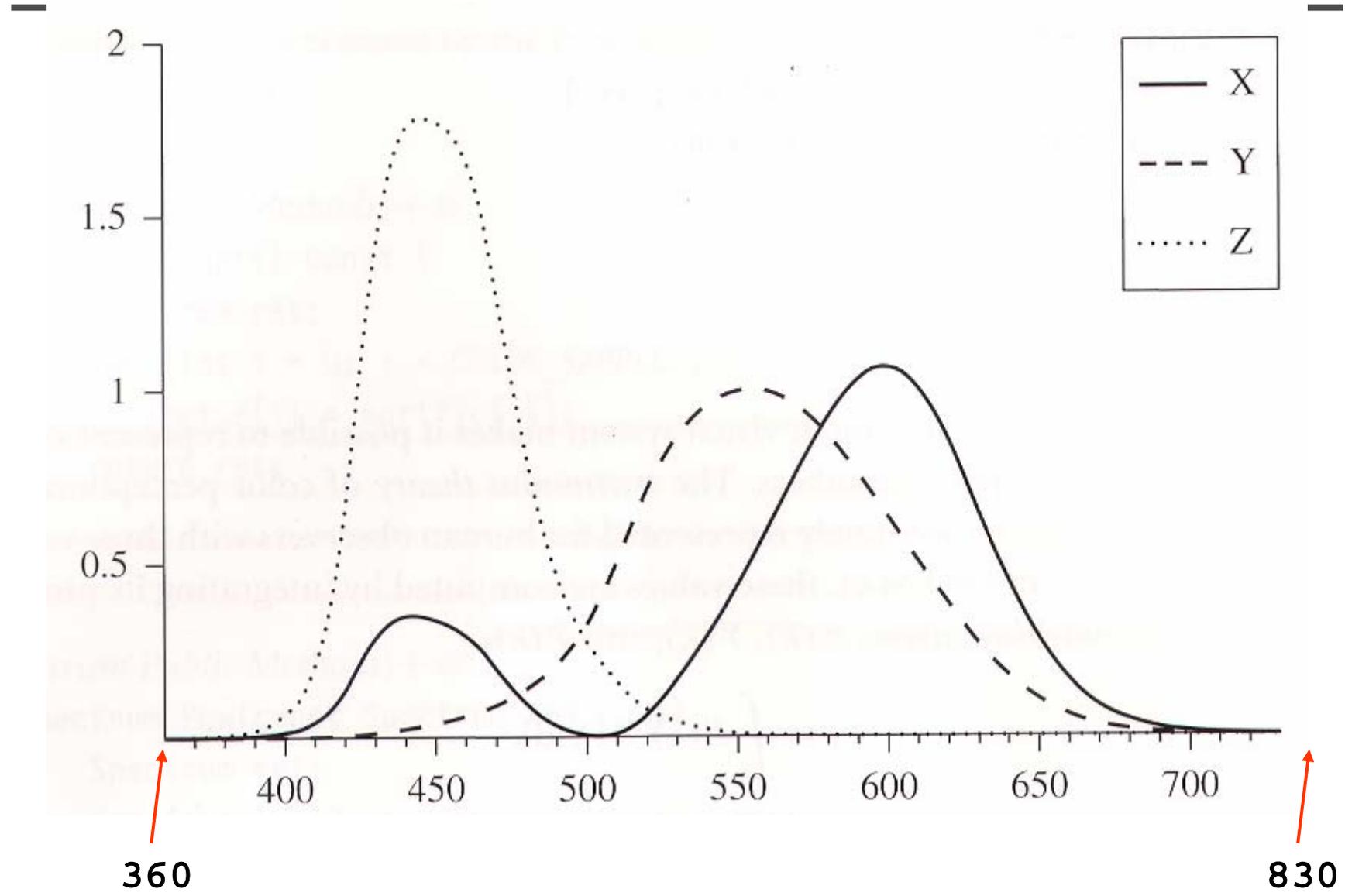


- Tristimulus theory: all visible SPDs can be accurately represented for human observers with three values, x_λ , y_λ and z_λ .
- The basis are the *spectral matching curves*, $X(\lambda)$, $Y(\lambda)$ and $Z(\lambda)$.
- Not a good model for computation

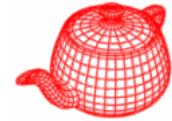
$$x_\lambda = \int_\lambda S(\lambda) X(\lambda) d\lambda$$

$$y_\lambda = \int_\lambda S(\lambda) Y(\lambda) d\lambda$$

$$z_\lambda = \int_\lambda S(\lambda) Z(\lambda) d\lambda$$

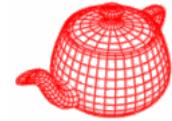


XYZ color



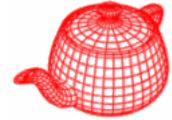
```
void XYZ(float xyz[3]) const {
    xyz[0] = xyz[1] = xyz[2] = 0.;
    for (int i = 0; i < COLOR_SAMPLES; ++i) {
        xyz[0] += XWeight[i] * c[i];
        xyz[1] += YWeight[i] * c[i];
        xyz[2] += ZWeight[i] * c[i];
    }
}
```

pbrrt uses RGB



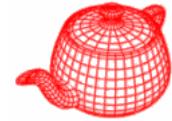
```
float Spectrum::XWeight[COLOR_SAMPLES] = {
    0.412453f, 0.357580f, 0.180423f
};
float Spectrum::YWeight[COLOR_SAMPLES] = {
    0.212671f, 0.715160f, 0.072169f
};
float Spectrum::ZWeight[COLOR_SAMPLES] = {
    0.019334f, 0.119193f, 0.950227f
};
Spectrum FromXYZ(float x, float y, float z) {
    float c[3];
    c[0] = 3.240479f * x + -1.537150f * y + -
0.498535f * z;
    c[1] = -0.969256f * x + 1.875991f * y +
0.041556f * z;
    c[2] = 0.055648f * x + -0.204043f * y +
1.057311f * z;
    return Spectrum(c);
}
```

Basic radiometry

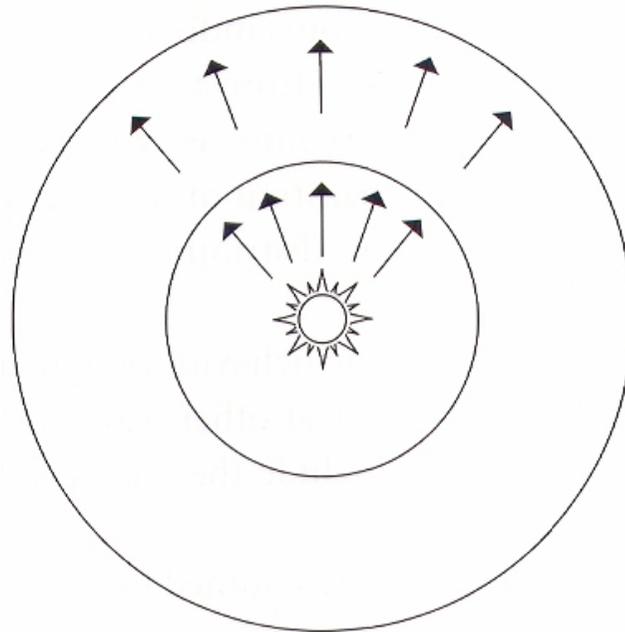


- Pbrt is based on Radiative transfer: study of the transfer of radiant energy based on radiometric principles and operates at the geometric optics level (light interacts with objects much larger than the light's wavelength)
- Based on the particle model, diffraction and interference can't be easily accounted for.

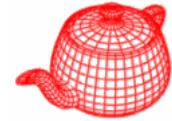
Flux (Φ)



- Radiant flux, power
- Total amount of energy passing through a surface per unit of time (J/s, W)

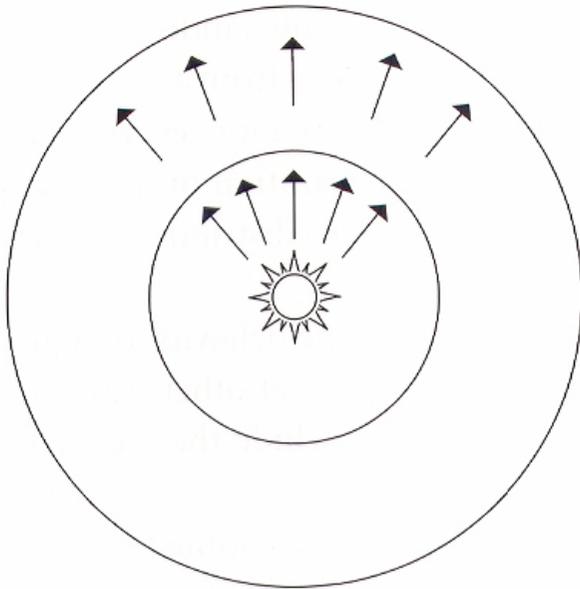


Irradiance (E)

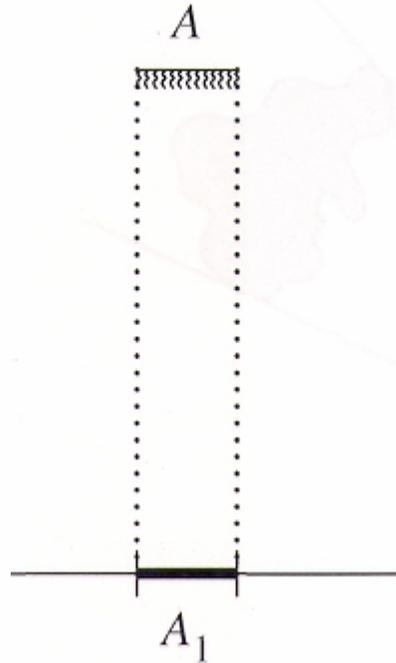


- Area density of flux (W/m^2) $E = \frac{d\Phi}{dA}$

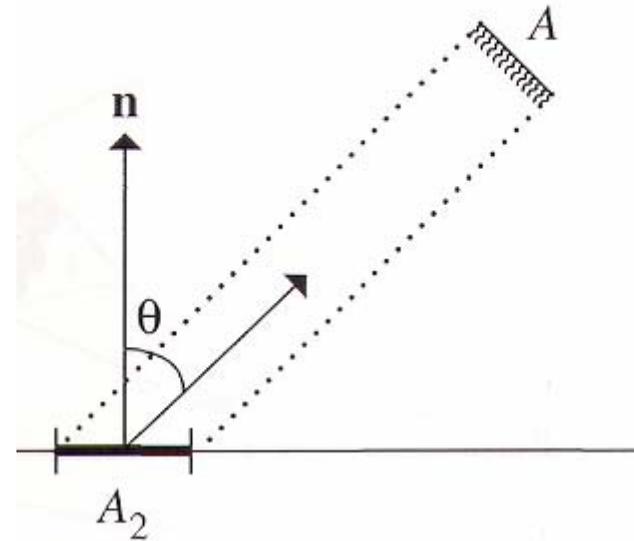
$$E = \frac{\Phi}{4\pi r^2}$$



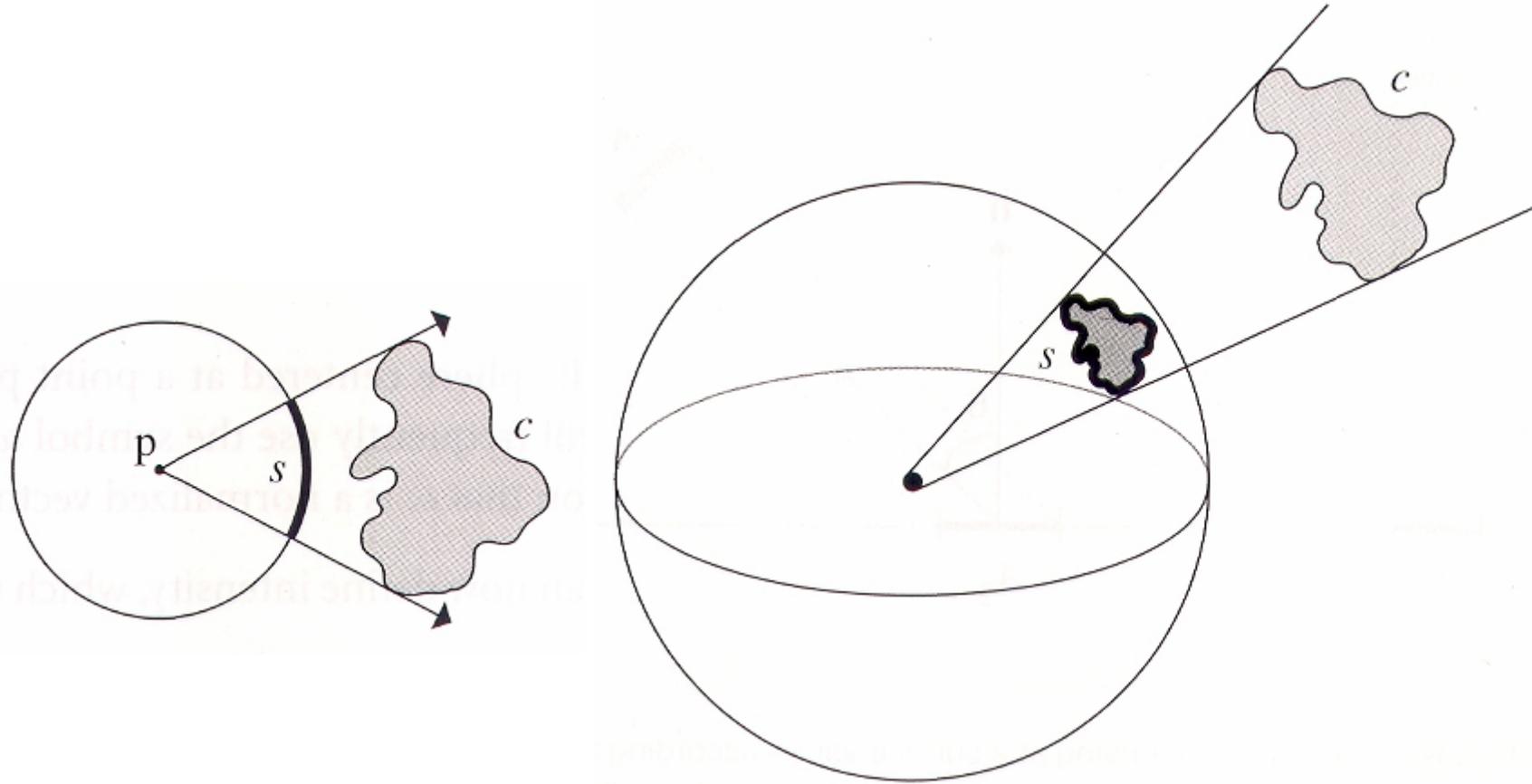
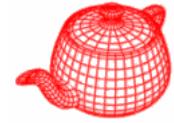
$$E = \frac{\Phi}{A}$$



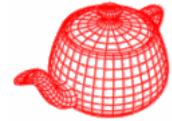
$$E = \frac{\Phi \cos \theta}{A}$$



Solid angle

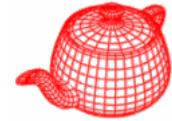


Intensity (I)



- Flux density per solid angle $I = \frac{d\Phi}{d\omega}$
- Intensity describes the directional distribution of light

Radiance (L)



- Flux density per unit area per solid angle

$$L = \frac{d\Phi}{d\omega dA^\perp}$$

- Most frequently used, constant along ray.
- All other quantities can be derived from radiance

