

Cameras

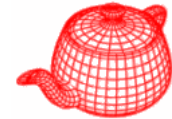
Digital Image Synthesis

Yung-Yu Chuang

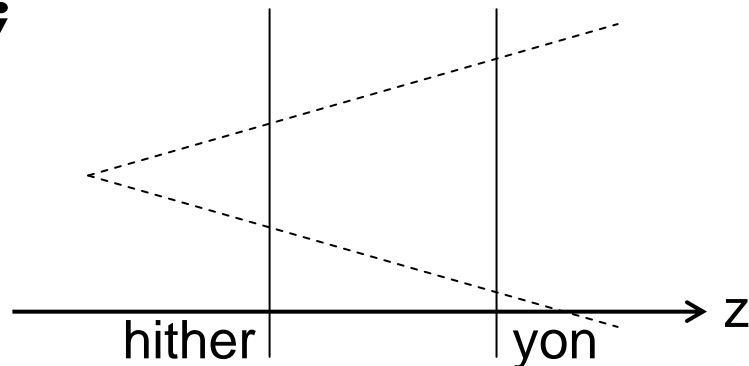
10/25/2007

with slides by Pat Hanrahan and Matt Pharr

Camera

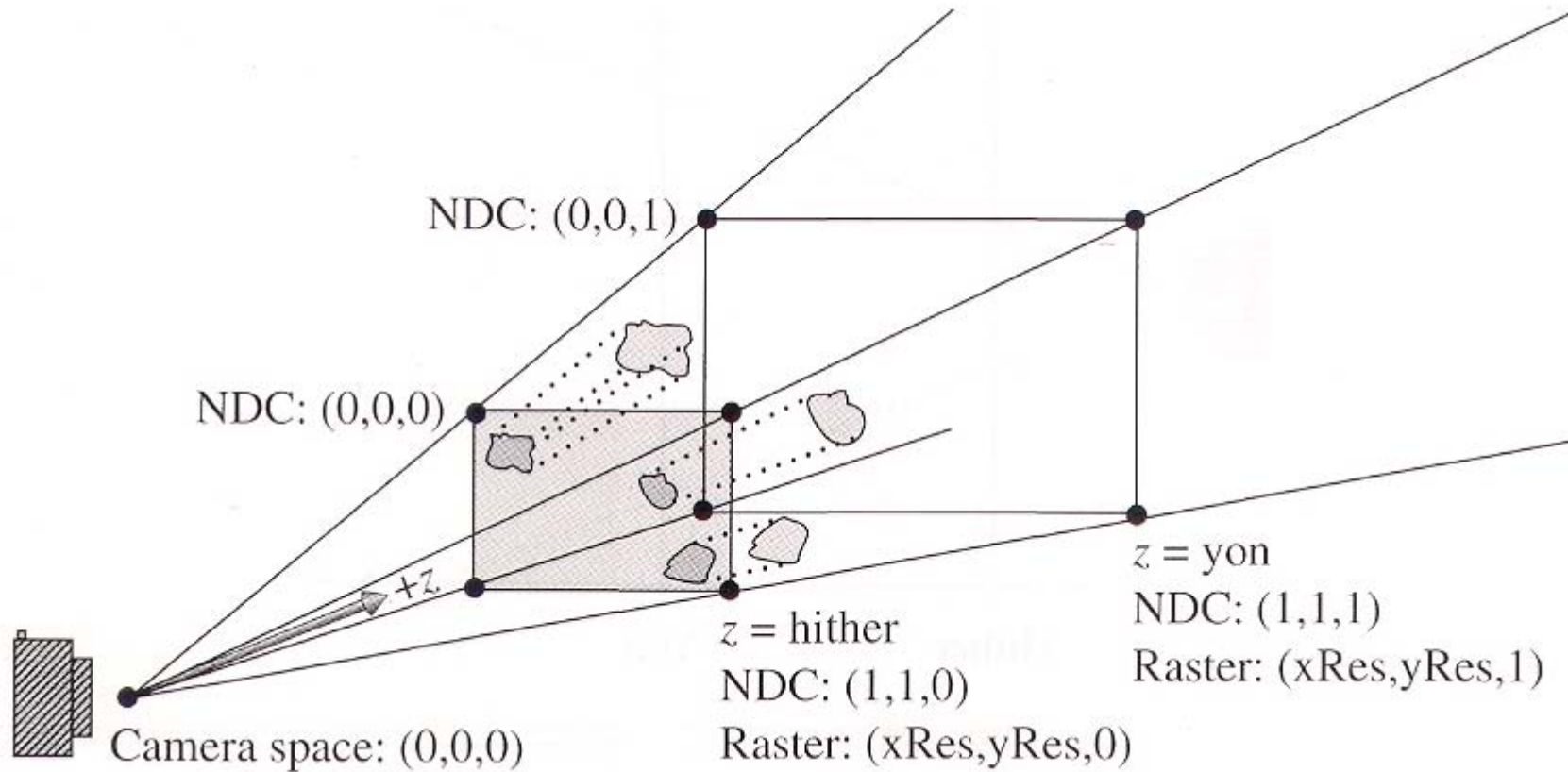
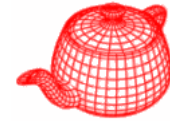


```
class Camera {  
public:    return a weight, useful for simulating real lens  
    virtual float GenerateRay(const Sample  
        &sample, Ray *ray) const = 0;  
    ...    sample position    corresponding  
    Film *film;    at the image plane    normalized ray in  
protected:    the world space  
    Transform WorldToCamera, CameraToWorld;  
    float ClipHither, ClipYon;  
    float ShutterOpen, ShutterClose;  
};
```

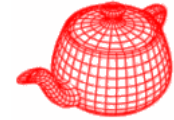


*for simulating
motion blur, not
Implemented yet*

Camera space

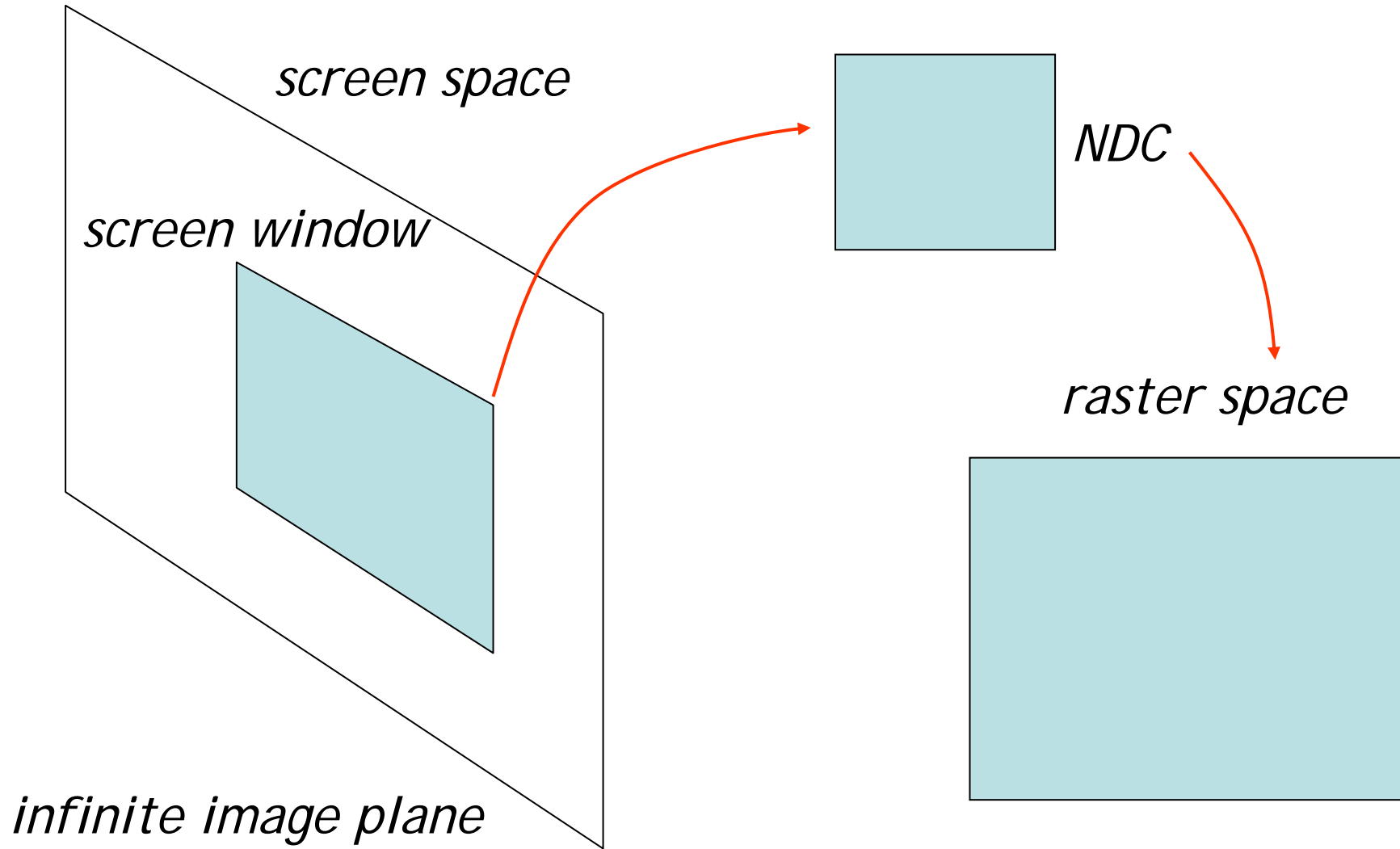
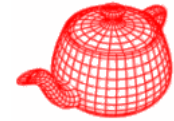


Coordinate spaces

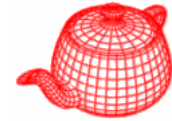


- world space
- object space
- camera space (origin: camera position, z: viewing direction, y: up direction)
- screen space: a 3D space defined on the image plane, z ranges from 0(near) to 1(far)
- normalized device space (NDC): (x, y) ranges from $(0,0)$ to $(1,1)$ for the rendered image, z is the same as the screen space
- raster space: similar to NDC, but the range of (x,y) is from $(0,0)$ to $(xRes, yRes)$

Screen space



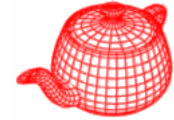
Projective camera models



- Transform a 3D scene coordinate to a 2D image coordinate by a 4x4 projective matrix

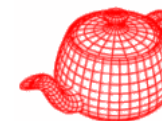
```
class ProjectiveCamera : public Camera {
public: camera to screen projection (3D to 2D)
    ProjectiveCamera(Transform &world2cam,
        Transform &proj, float Screen[4],
        float hither, float yon, float sopen,
        float sclose, float lensr, float focald,
        Film *film);
protected:
    Transform CameraToScreen, WorldToScreen,
        RasterToCamera;
    Transform ScreenToRaster, RasterToScreen;
    float LensRadius, FocalDistance;
};
```

Projective camera models

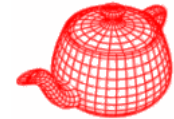


```
ProjectiveCamera::ProjectiveCamera(...)  
    :Camera(w2c, hither, yon, sopen, sclose, f) {  
    ...  
    CameraToScreen=proj;  
    WorldToScreen=CameraToScreen*WorldToCamera;  
    ScreenToRaster  
        = Scale(float(film->xResolution),  
                float(film->yResolution), 1.f)*  
          Scale(1.f / (Screen[1] - Screen[0]),  
                1.f / (Screen[2] - Screen[3]), 1.f)*  
          Translate(Vector(-Screen[0], -Screen[3], 0.f));  
    RasterToScreen = ScreenToRaster.GetInverse();  
    RasterToCamera =  
        CameraToScreen.GetInverse() * RasterToScreen;  
}
```

Projective camera models



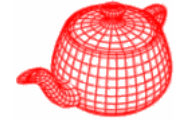
Orthographic camera



```
Transform Orthographic(float znear,  
                       float zfar)  
{  
    return Scale(1.f, 1.f, 1.f/(zfar-znear))  
        *Translate(Vector(0.f, 0.f, -znear));  
}
```

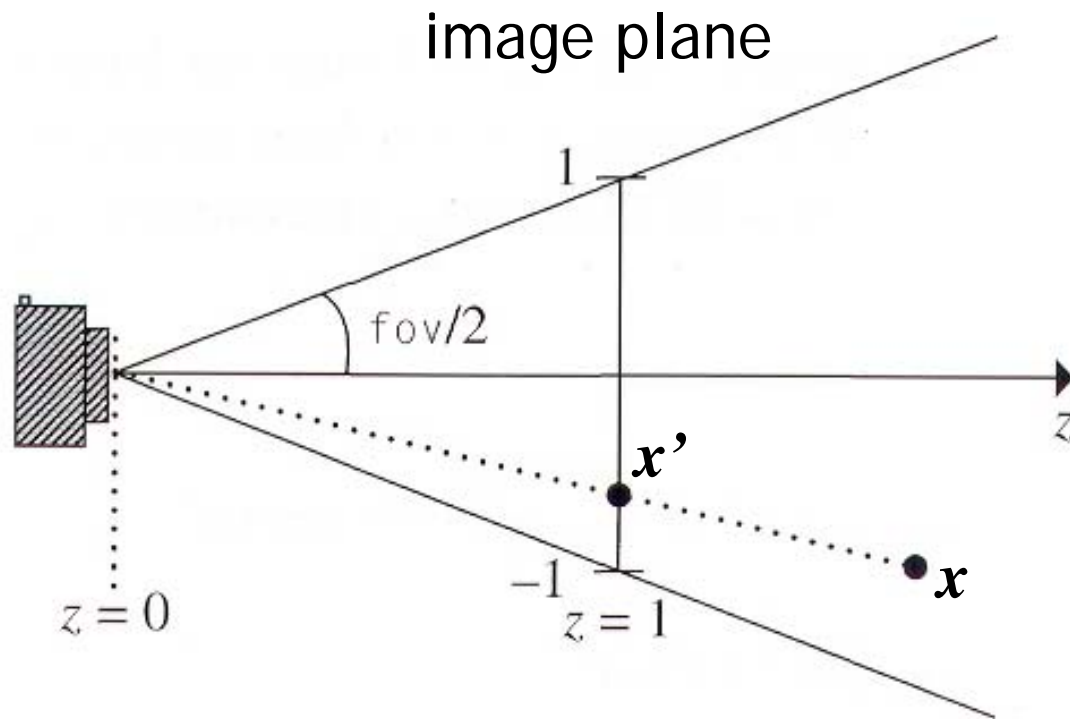
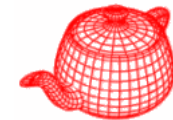
```
OrthoCamera::OrthoCamera( ... )  
    : ProjectiveCamera(world2cam,  
        Orthographic(hither, yon),  
        Screen, hither, yon, sopen, sclose,  
        lensr, focald, f) {  
}
```

OrthoCamera::GenerateRay



```
float OrthoCamera::GenerateRay
    (const Sample &sample, Ray *ray) const {
    Point Pras(sample.imageX, sample.imageY, 0);
    Point Pcamera;
    RasterToCamera(Pras, &Pcamera);
    ray->o = Pcamera;
    ray->d = Vector(0,0,1);
    <Modify ray for depth of field>
    ray->mint = 0.;
    ray->maxt = ClipYon - ClipHither;
    ray->d = Normalize(ray->d);
    CameraToWorld(*ray, ray);
    return 1.f;
}
```

Perspective camera

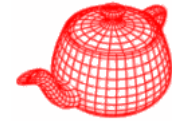


$$\begin{matrix}
 x' = x / z \\
 y' = y / z
 \end{matrix}
 \begin{bmatrix}
 1 & 0 & 0 & 0 \\
 0 & 1 & 0 & 0 \\
 ? & ? & ? & ? \\
 0 & 0 & 1 & 0
 \end{bmatrix}
 \begin{bmatrix}
 x \\
 y \\
 z \\
 1
 \end{bmatrix}$$

$$z' = \frac{f(z-n)}{z(f-n)}$$

f, n, z?

Perspective camera



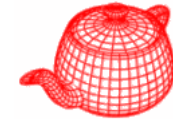
```
Transform Perspective(float fov,float n,float f)
{
    near_z    far_z
    float inv_denom = 1.f/(f-n);
    Matrix4x4 *persp =
    new Matrix4x4(1, 0,      0,      0,
                  0, 1,      0,      0,
                  0, 0,  f*inv_denom, -f*n*inv_denom,
                  0, 0,      1,      0);

    float invTanAng= 1.f / tanf(Radians(fov)/2.f);
    return Scale(invTanAng, invTanAng, 1) *
           Transform(persp);
}
```

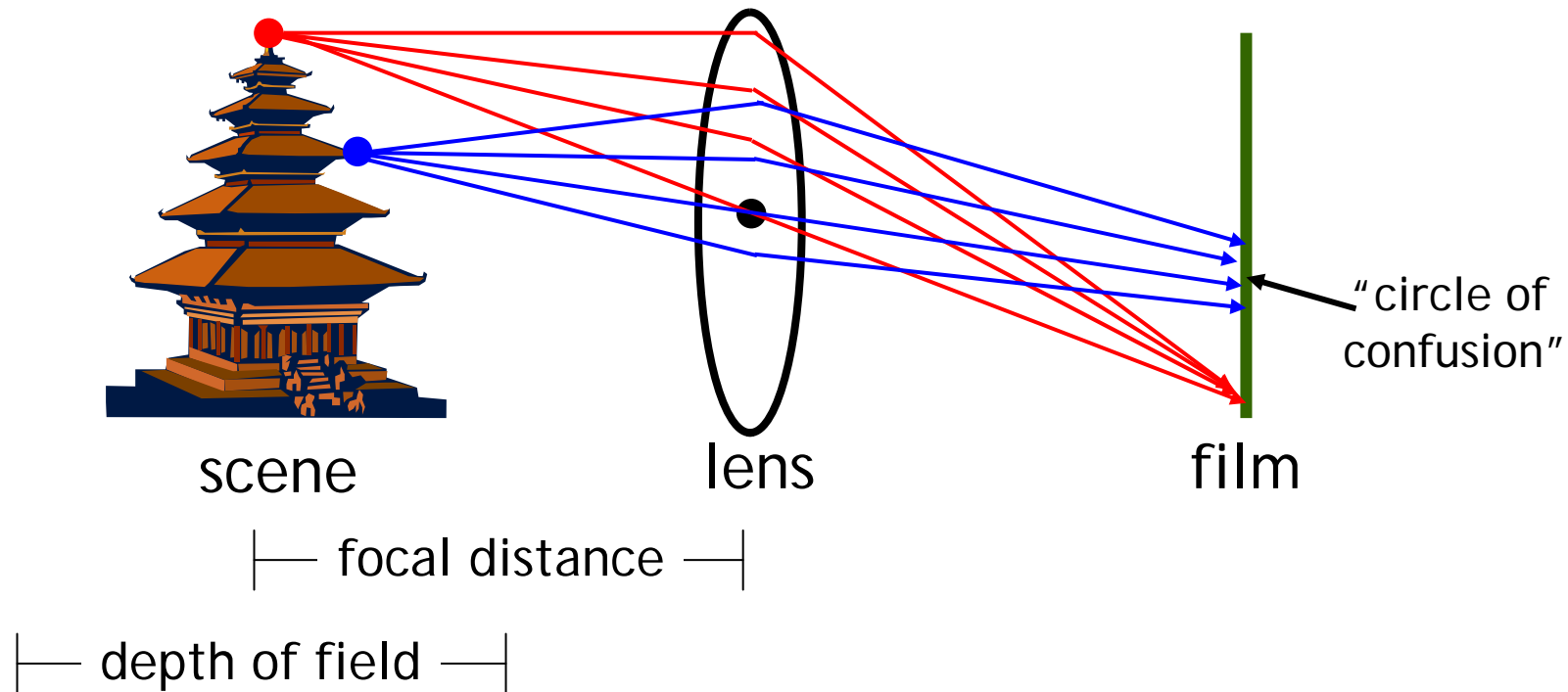
PerspectiveCamera::GenerateRay

```
float PerspectiveCamera::GenerateRay
    (const Sample &sample, Ray *ray) const
{
    // Generate raster and camera samples
    Point Pras(sample.imageX, sample.imageY, 0);
    Point Pcamera;
    RasterToCamera(Pras, &Pcamera);
    ray->o = Pcamera;
    ray->d = Vector(Pcamera.x, Pcamera.y, Pcamera.z);
    <Modify ray for depth of field>
    ray->d = Normalize(ray->d);
    ray->mint = 0.;
    ray->maxt = (ClipYon-ClipHither)/ray->d.z;
    CameraToWorld(*ray, ray);
    return 1.f;
}
```

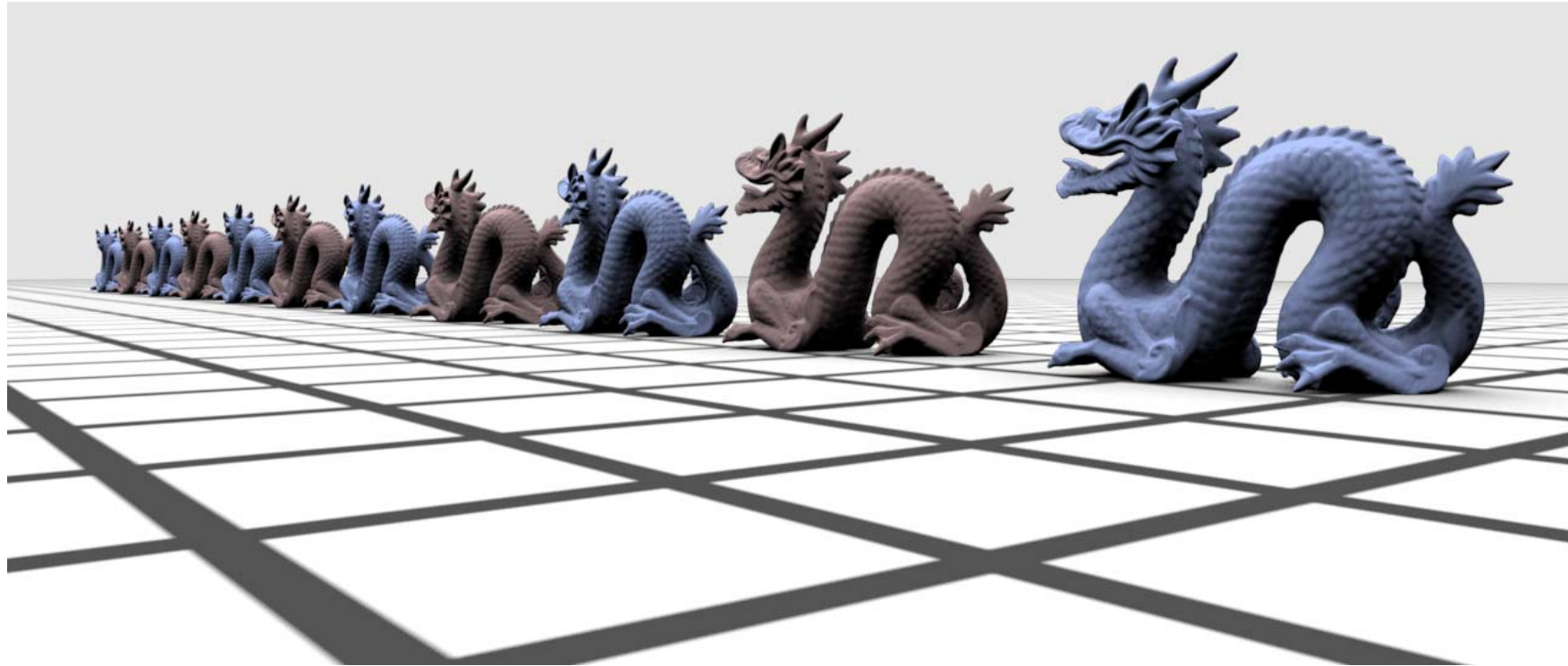
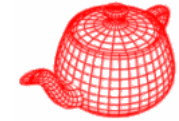
Depth of field



- Circle of confusion $\frac{1}{d_o} + \frac{1}{d_i} = \frac{1}{f}$
- Depth of field: the range of distances from the lens at which objects appear in focus (circle of confusion roughly smaller than a pixel)

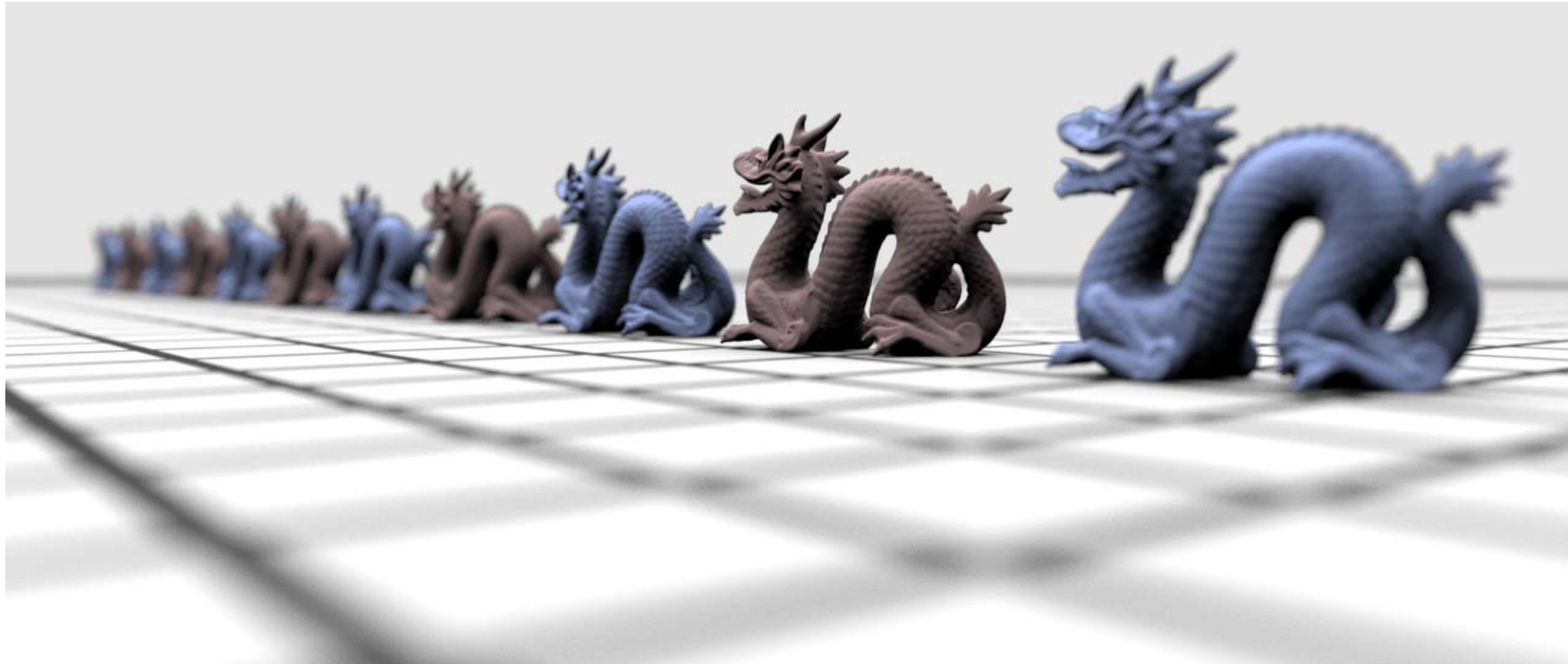
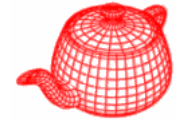


Depth of field



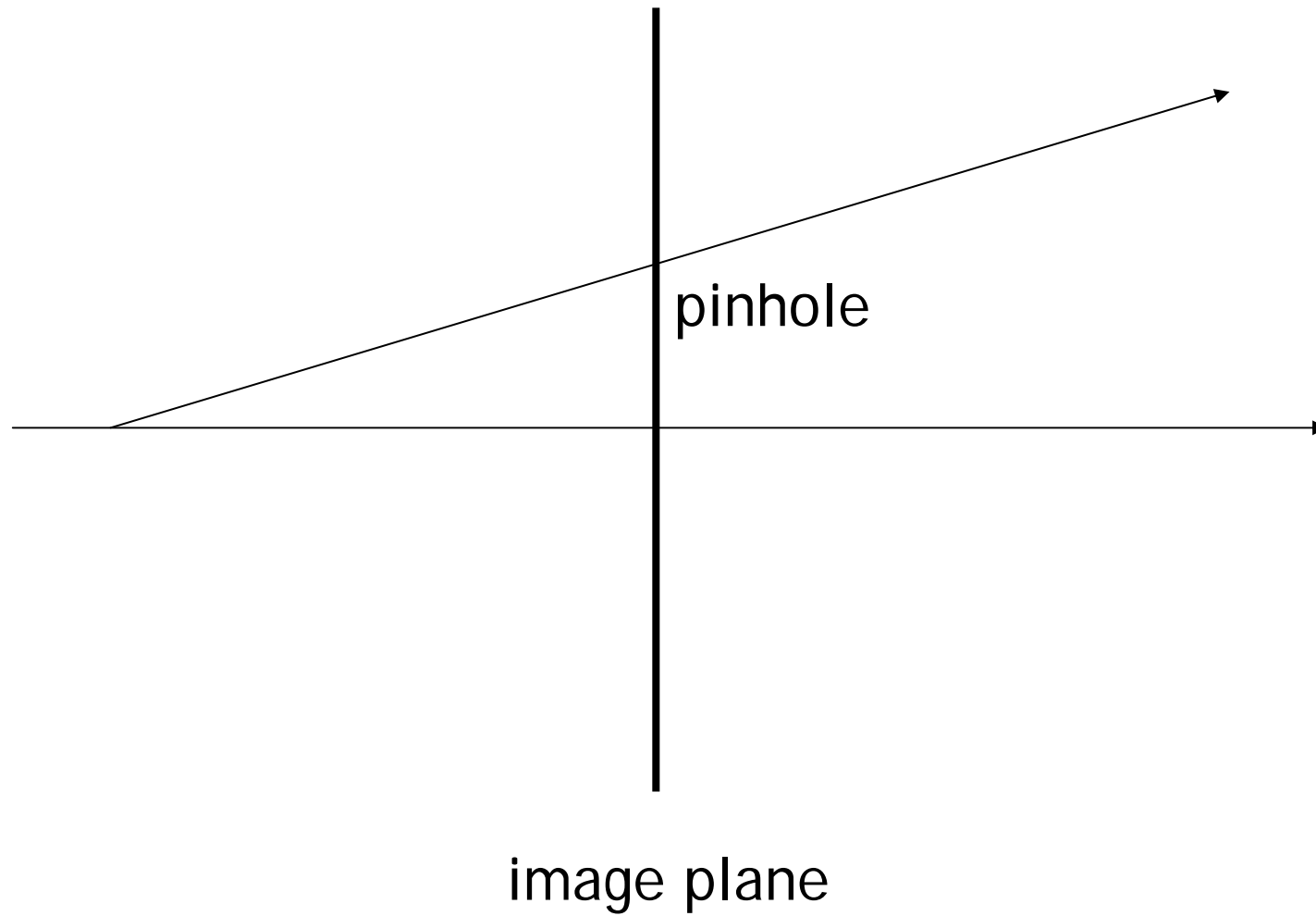
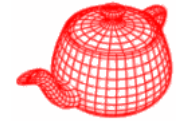
without depth of field

Depth of field

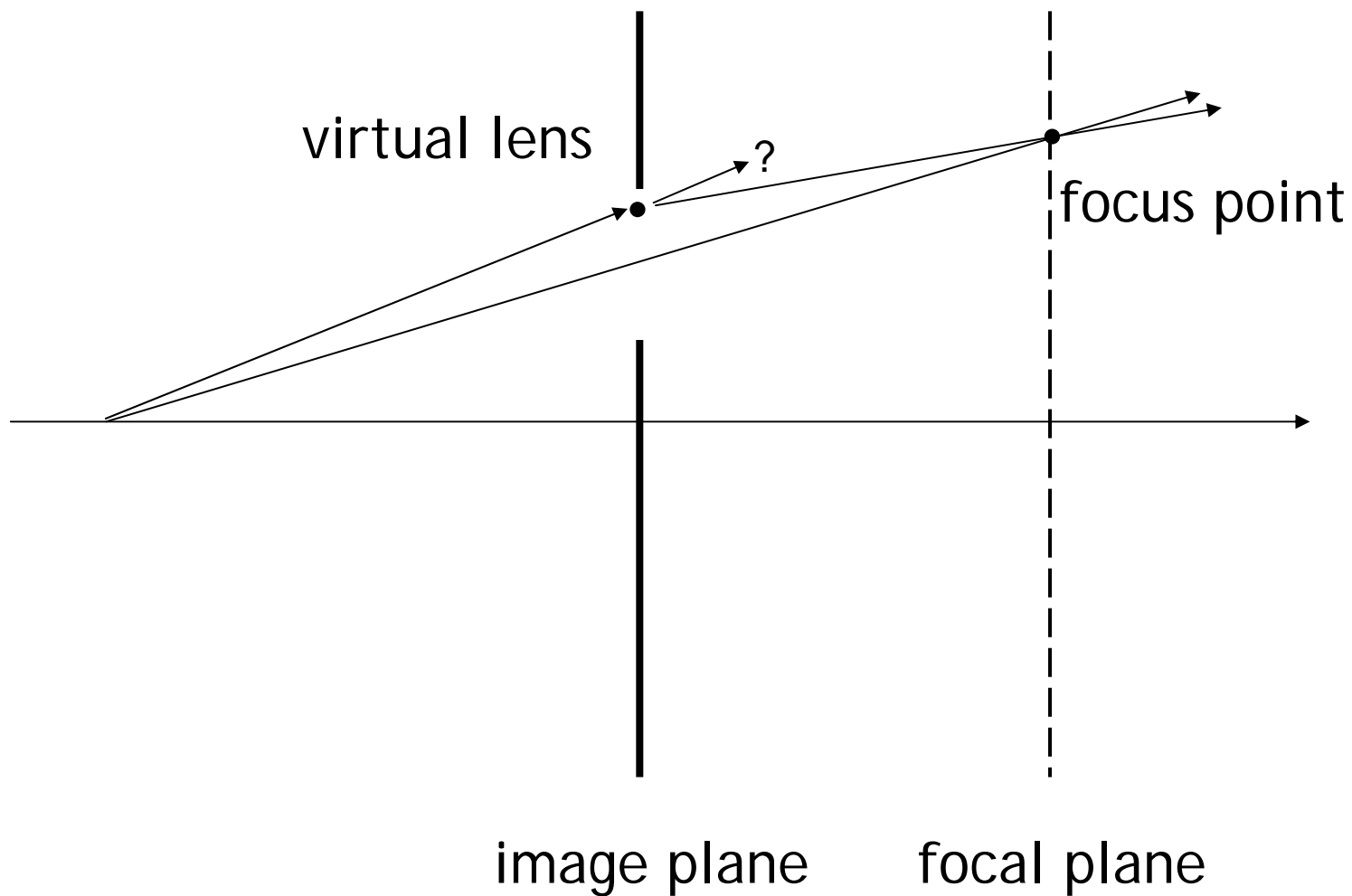
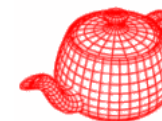


with depth of field

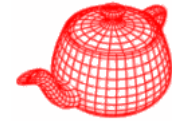
Sample the lens



Sample the lens



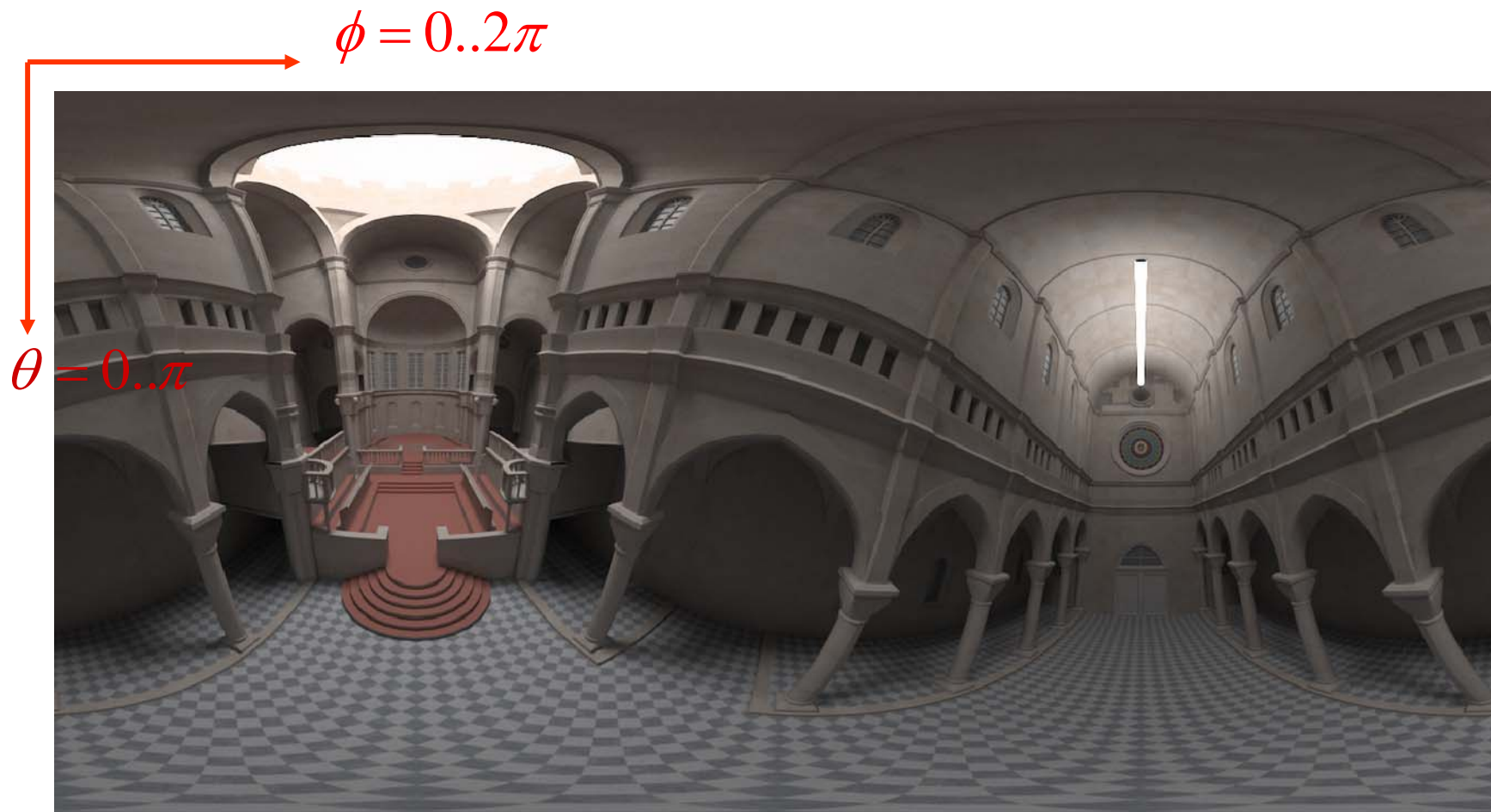
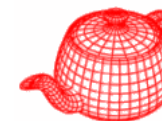
In GenerateRay(...)



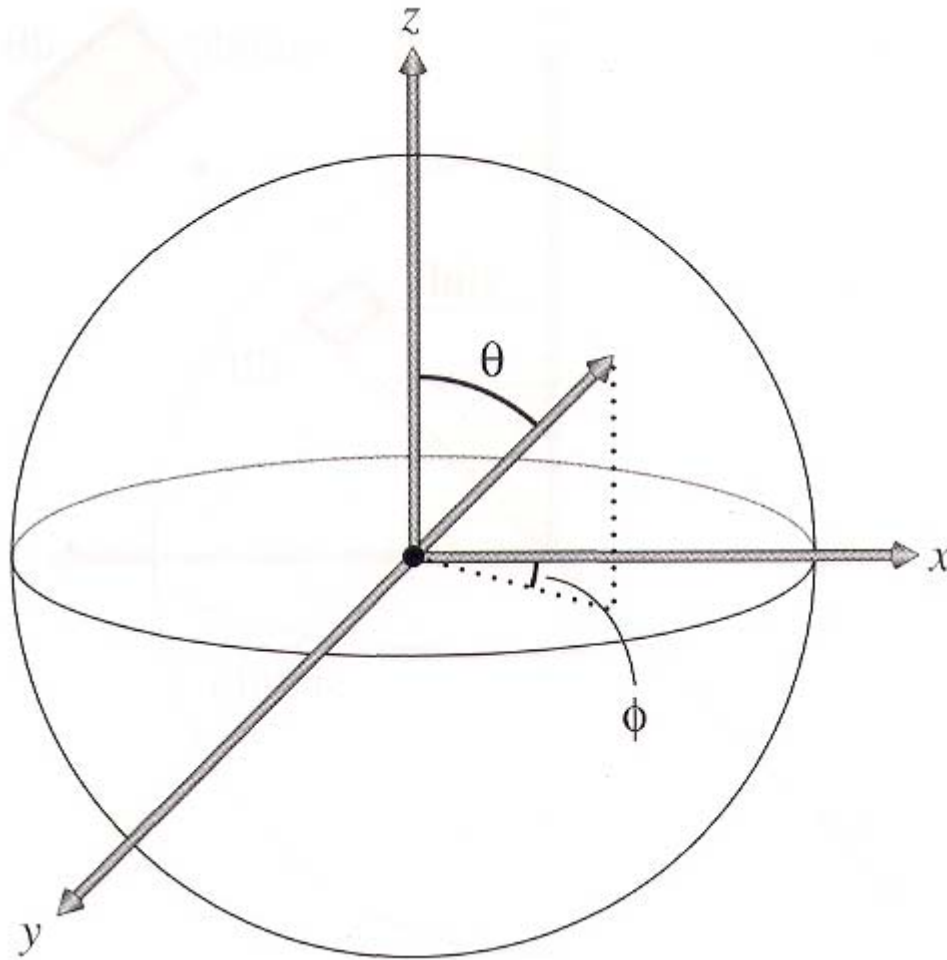
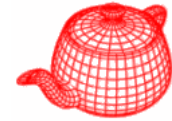
```
if (LensRadius > 0.) {
    // Sample point on lens
    float lensU, lensV;
    ConcentricSampleDisk(sample.lensU, sample.lensV,
                        &lensU, &lensV);

    lensU *= LensRadius;
    lensV *= LensRadius;
    // Compute point on plane of focus
    float ft = (FocalDistance - ClipHither) / ray->d.z;
    Point Pfocus = (*ray)(ft);
    // Update ray for effect of lens
    ray->o.x += lensU;
    ray->o.y += lensV;
    ray->d = Pfocus - ray->o;
}
```

Environment camera

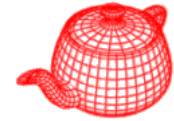


Environment camera



$$\begin{aligned}x &= \sin \theta \cos \phi \\y &= \sin \theta \sin \phi \\z &= \cos \theta\end{aligned}$$

EnvironmentCamera



```
EnvironmentCamera::
```

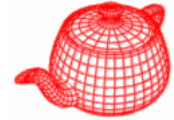
```
    EnvironmentCamera(const Transform &world2cam,  
                      float hither, float yon,  
                      float sopen, float sclose,  
                      Film *film)
```

```
    : Camera(world2cam, hither, yon,  
             sopen, sclose, film)
```

```
{  
    rayOrigin = CameraToWorld(Point(0,0,0));  
}
```

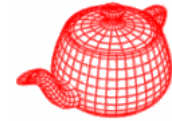
↑
in world space

EnvironmentCamera::GenerateRay

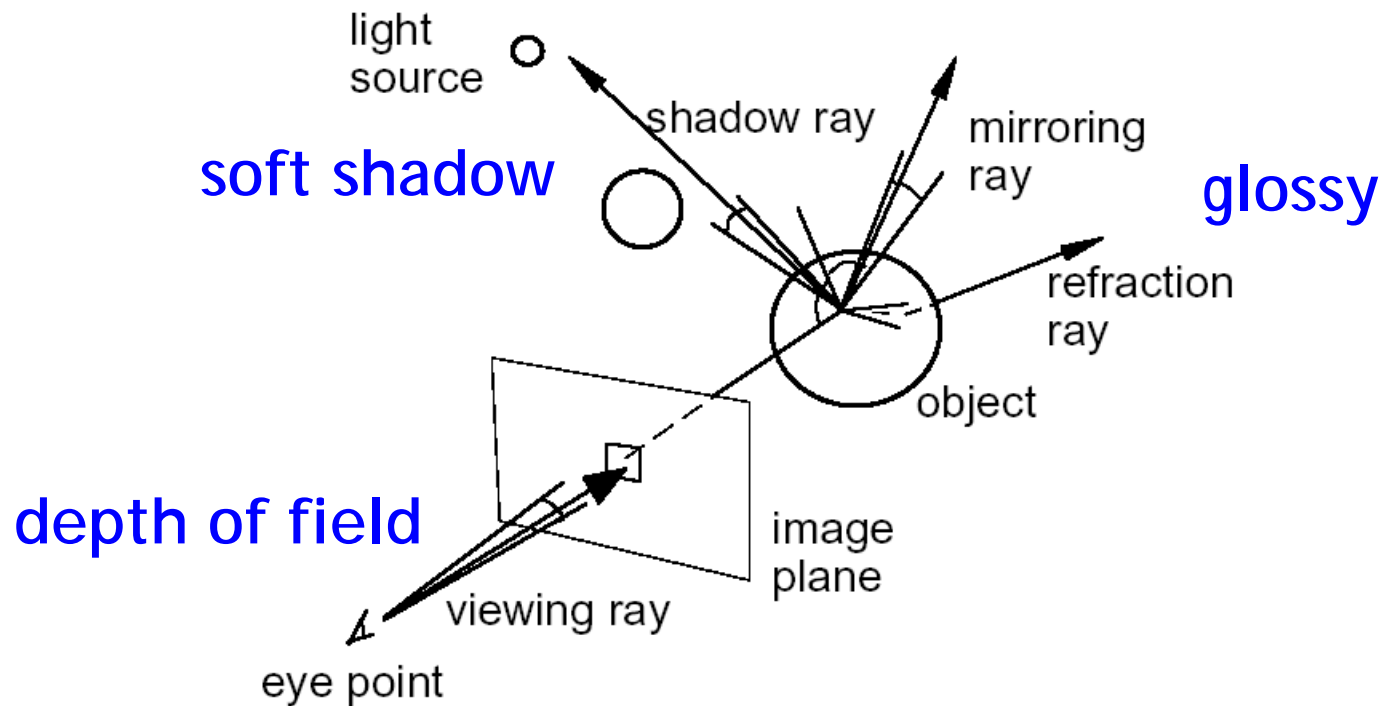


```
float EnvironmentCamera::GenerateRay
    (const Sample &sample, Ray *ray) const
{
    ray->o = rayOrigin;
    float theta=M_PI*sample.imageY/film->yResolution;
    float phi=2*M_PI*sample.imageX/film->xResolution;
    Vector dir(sinf(theta)*cosf(phi), cosf(theta),
               sinf(theta)*sinf(phi));
    CameraToWorld(dir, &ray->d);
    ray->mint = ClipHither;
    ray->maxt = ClipYon;
    return 1.f;
}
```

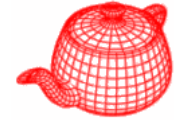
Distributed ray tracing



- *SIGGRAPH 1984*, by Robert L. Cook, Thomas Porter and Loren Carpenter from LucasFilm.
- Apply distribution-based sampling to many parts of the ray-tracing algorithm.



Distributed ray tracing



Gloss/Translucency

- Perturb directions reflection/transmission, with distribution based on angle from ideal ray

Depth of field

- Perturb eye position on lens

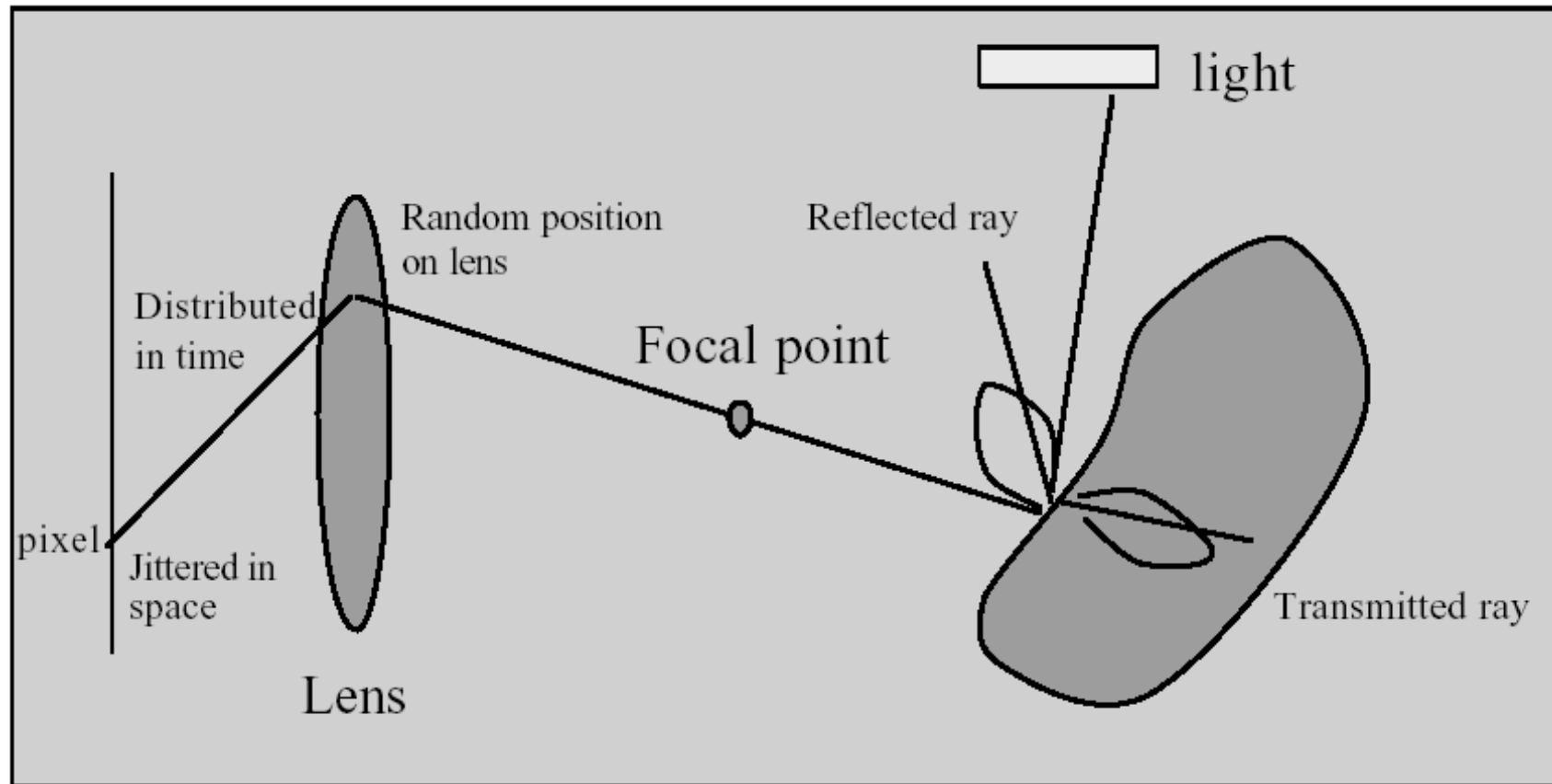
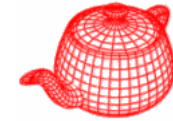
Soft shadow

- Perturb illumination rays across area light

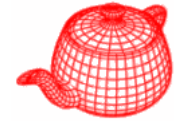
Motion blur

- Perturb eye ray samples in time

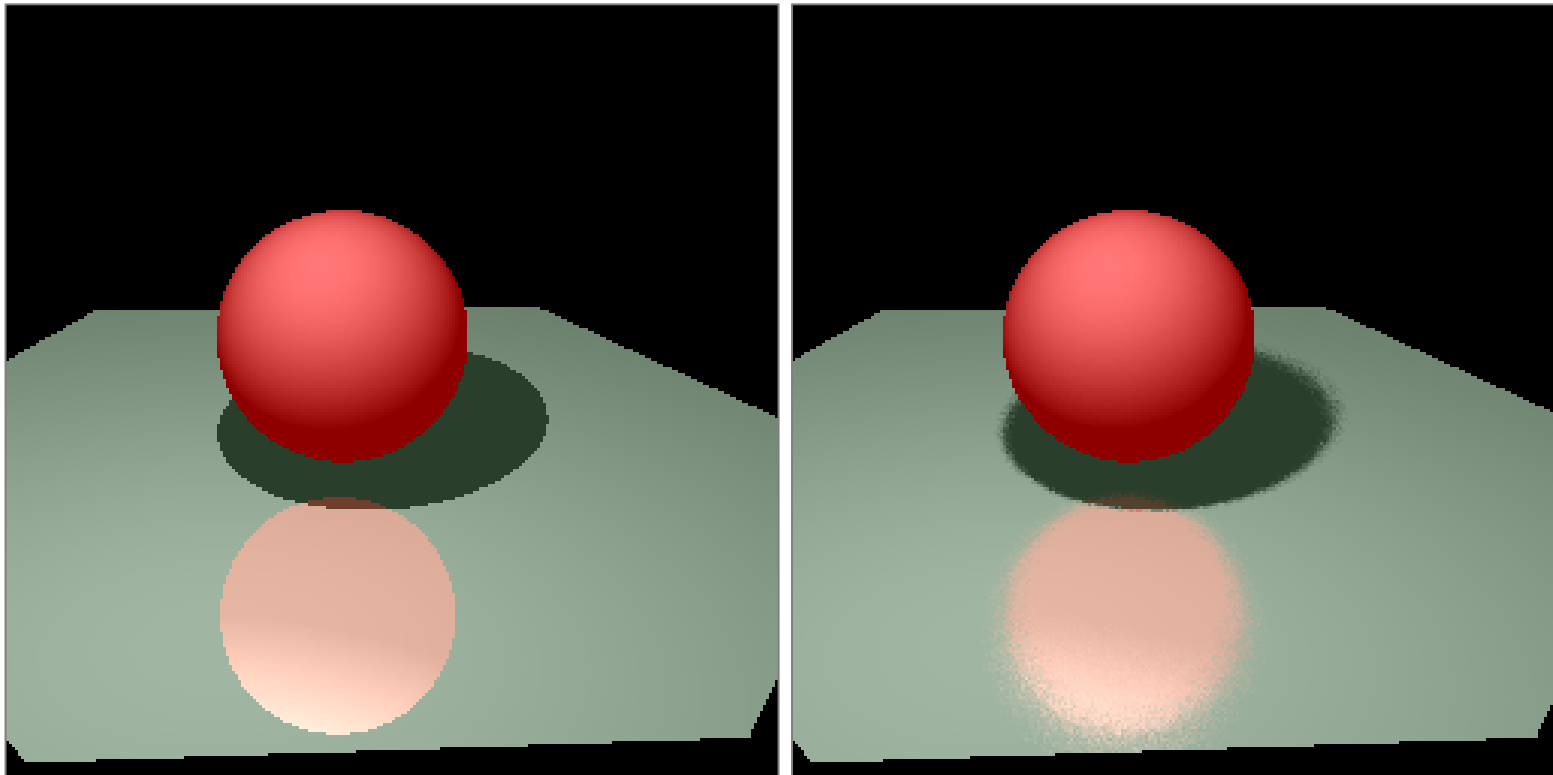
Distributed ray tracing



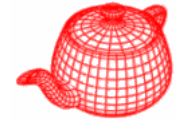
DRT: Gloss/Translucency



- Blurry reflections and refractions are produced by randomly perturbing the reflection and refraction rays from their "true" directions.



Glossy reflection

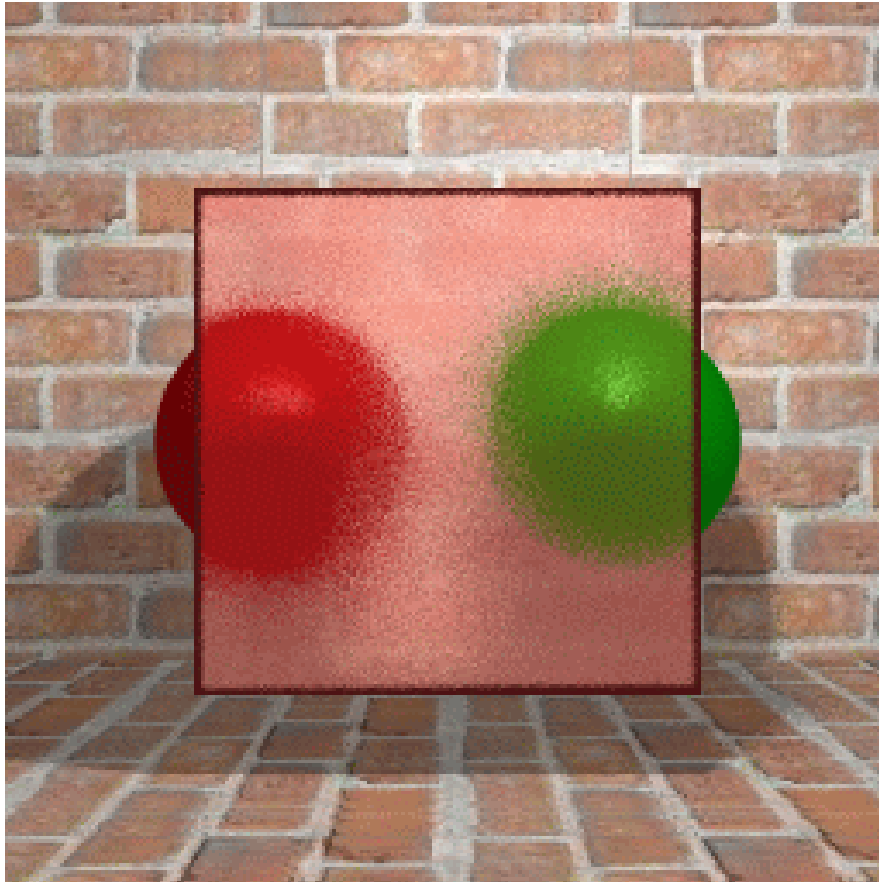
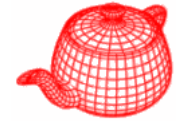


4 rays

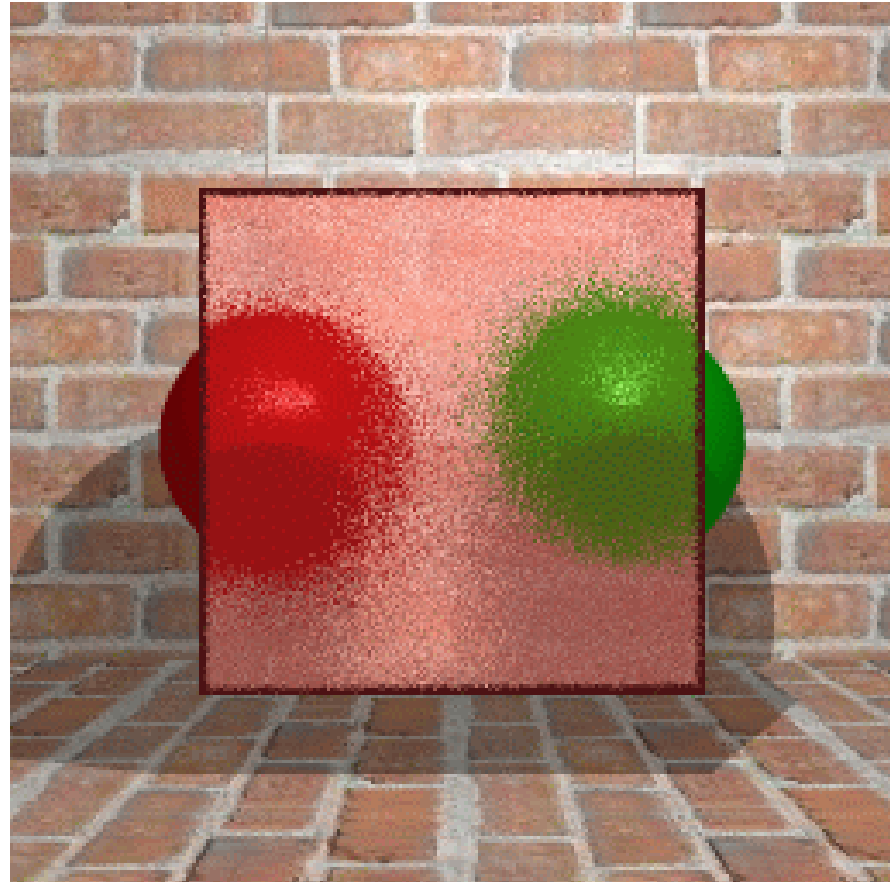


64 rays

Translucency

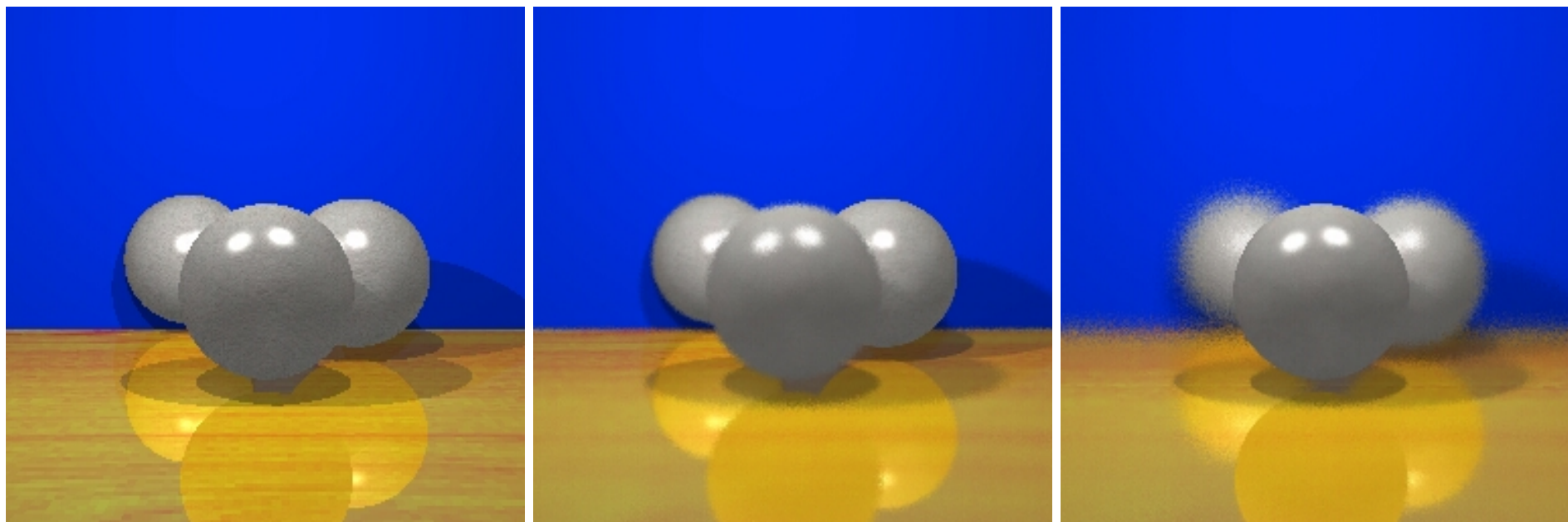
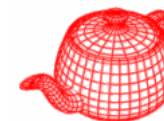


4 rays

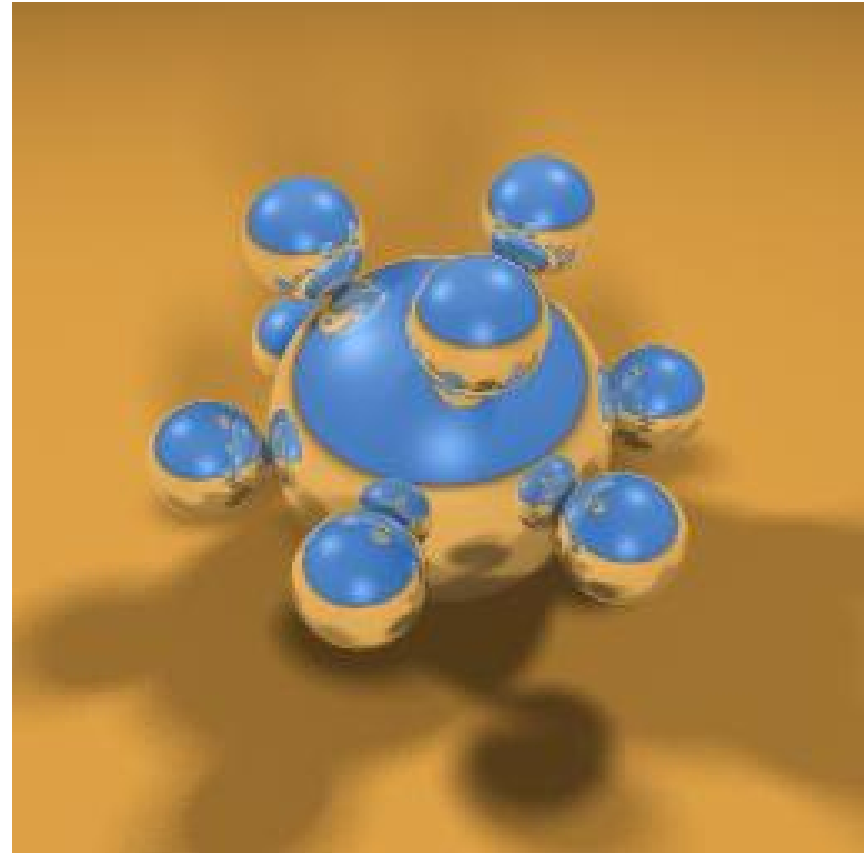
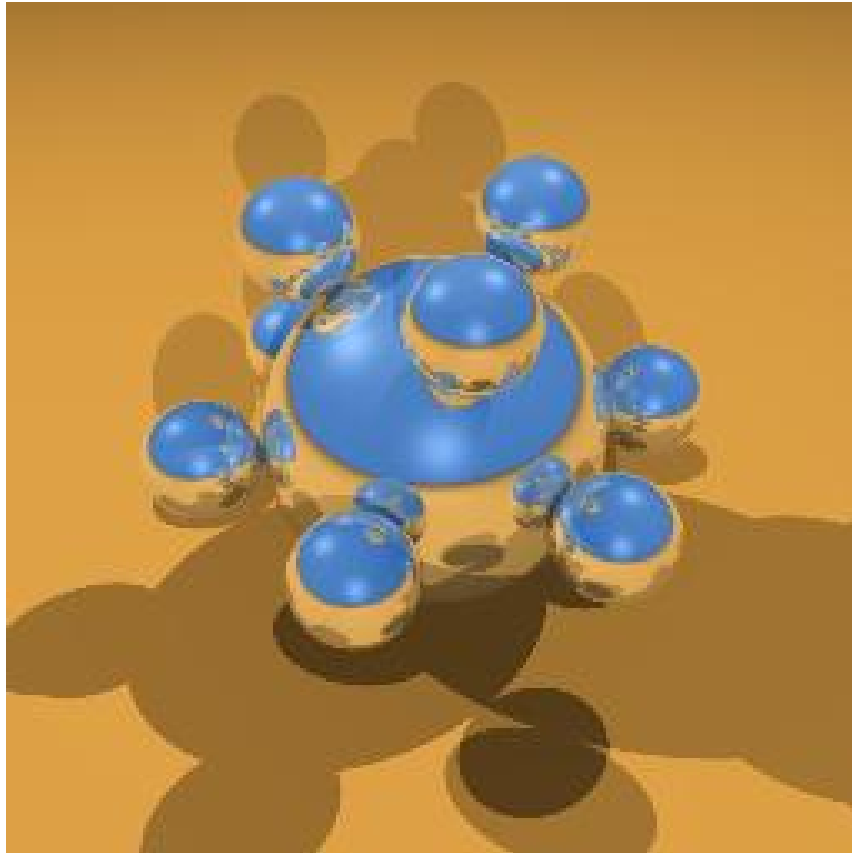
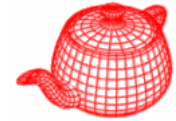


16 rays

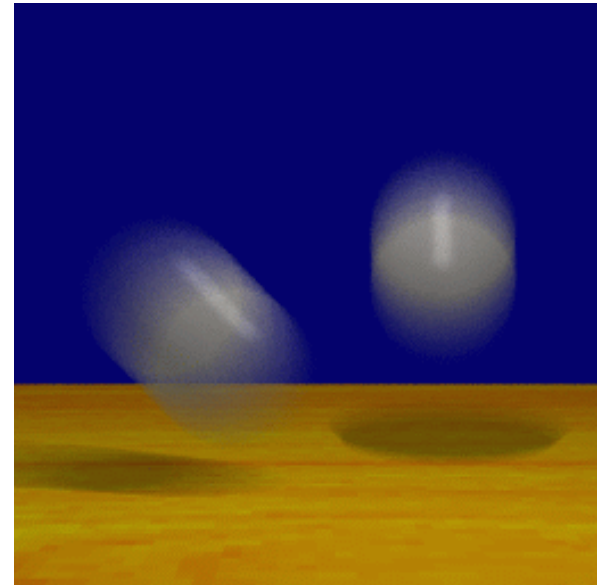
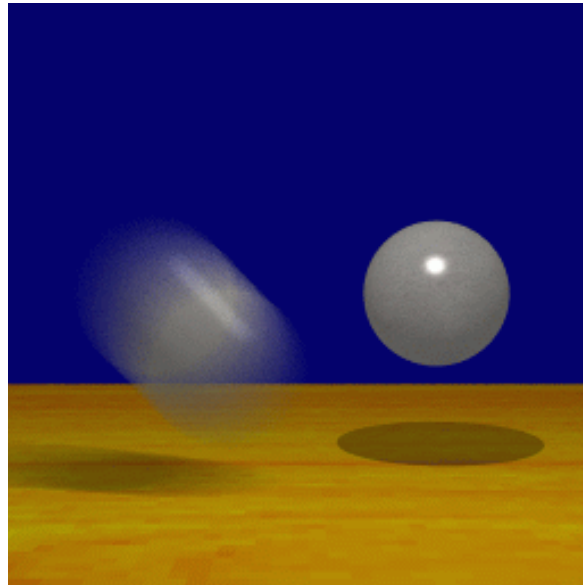
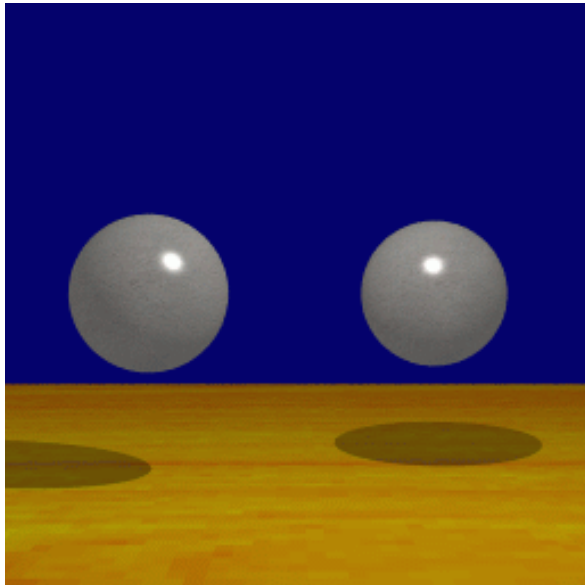
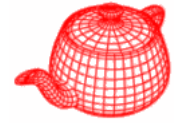
Depth of field



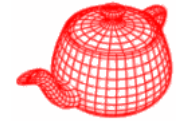
Soft shadows



Motion blur



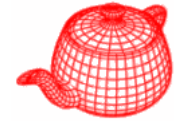
Results



Adventures of Andre & Wally B (1986)

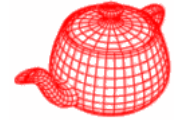


Realistic camera model



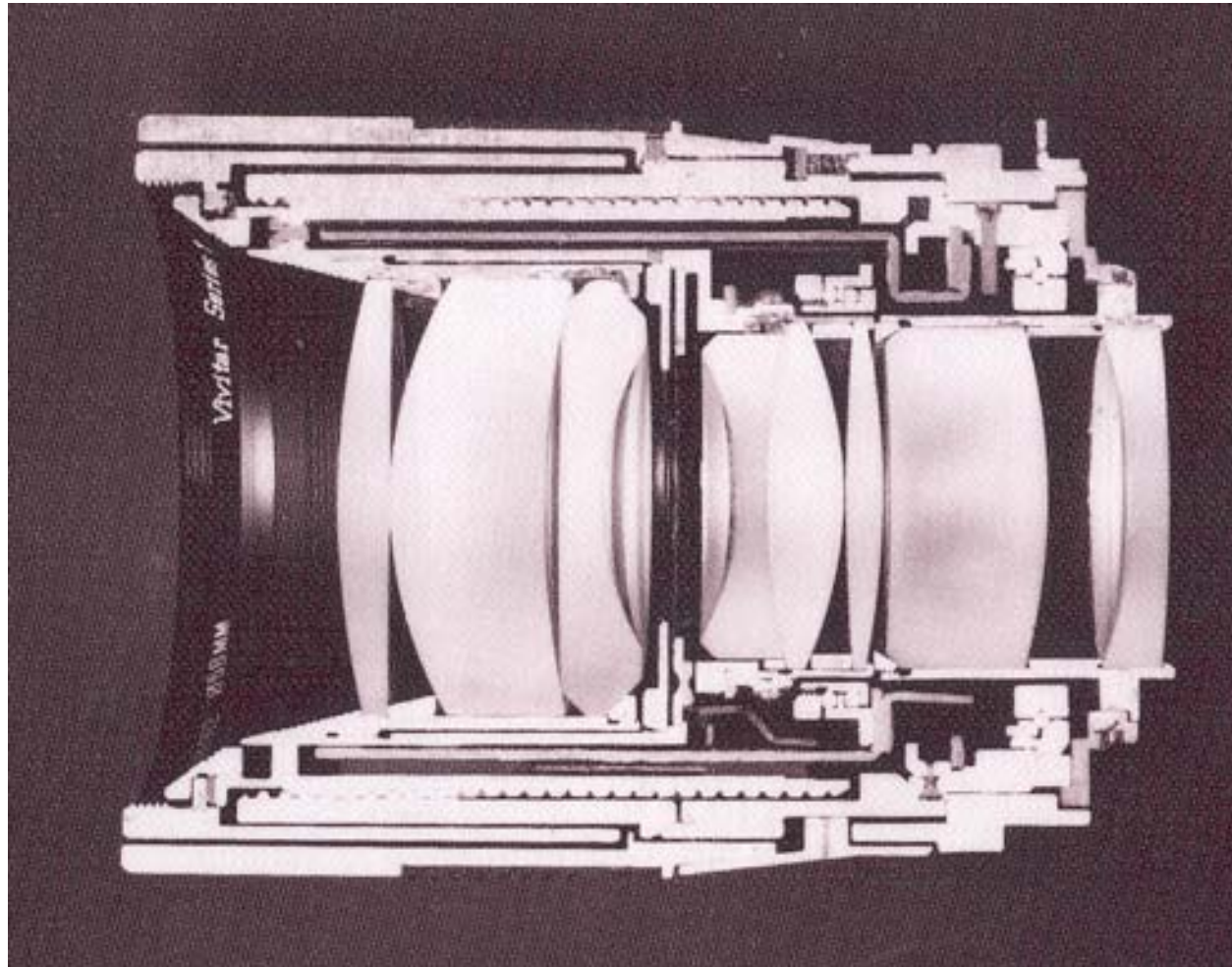
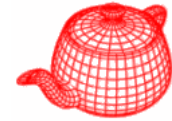
- Most camera models in graphics are not geometrically or radiometrically correct.
- Model a camera with a lens system and a film backplane. A lens system consists of a sequence of simple lens elements, stops and apertures.

Why a realistic camera model?



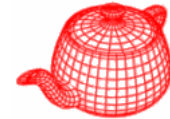
- Physically-based rendering. For more accurate comparison to empirical data.
- Seamlessly merge CGI and real scene, for example, VFX.
- For vision and scientific applications.
- The camera metaphor is familiar to most 3d graphics system users.

Real Lens



Cutaway section of a Vivitar Series 1 90mm f/2.5 lens
Cover photo, Kingslake, *Optics in Photography*

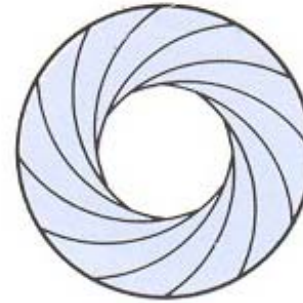
Exposure



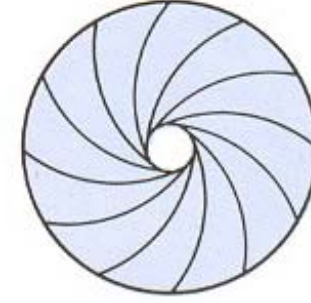
- Two main parameters:
 - Aperture (in f stop)



Full aperture



Medium aperture

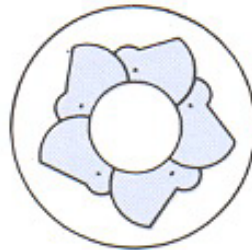


Stopped down

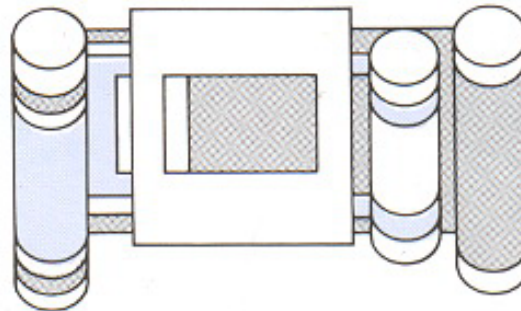
- Shutter speed (in fraction of a second)



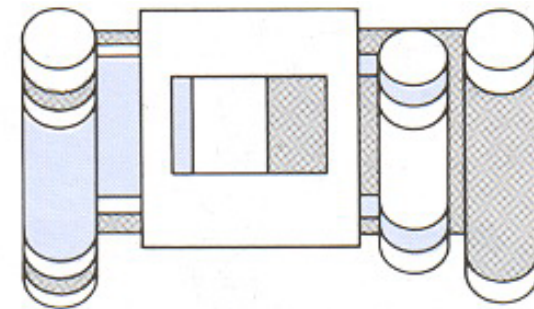
Blade (closing)



Blade (open)

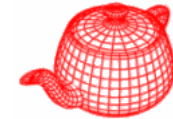


Focal plane (closed)



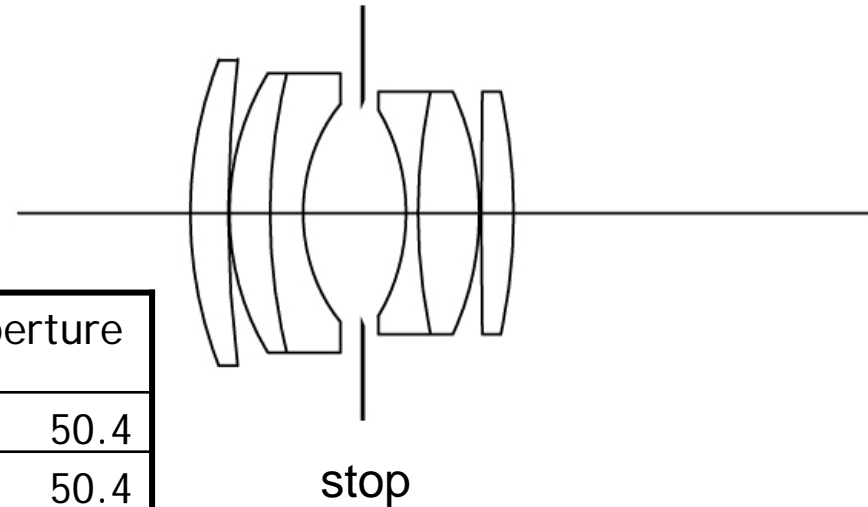
Focal plane (open)

Double Gauss

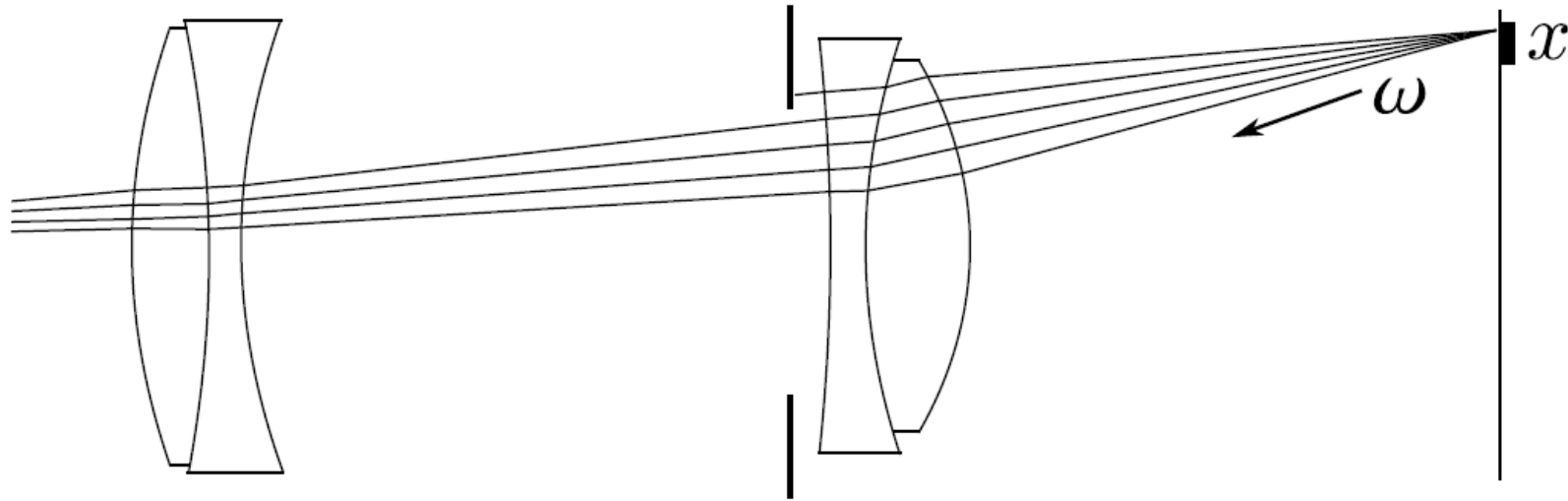
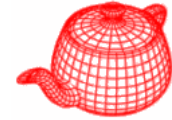


**Data from W. Smith,
Modern Lens Design, p 312**

Radius (mm)	Thick (mm)	n_d	V-no	aperture
58.950	7.520	1.670	47.1	50.4
169.660	0.240			50.4
38.550	8.050	1.670	47.1	46.0
81.540	6.550	1.699	30.1	46.0
25.500	11.410			36.0
	9.000			34.2
-28.990	2.360	1.603	38.0	34.0
81.540	12.130	1.658	57.3	40.0
-40.770	0.380			40.0
874.130	6.440	1.717	48.0	40.0
-79.460	72.228			40.0



Measurement equation



$$R = \int \int \int \int L(T(x, \omega, \lambda); \lambda) S(x, t) P(x, \lambda) \cos \theta \, dx \, d\omega \, dt \, d\lambda$$

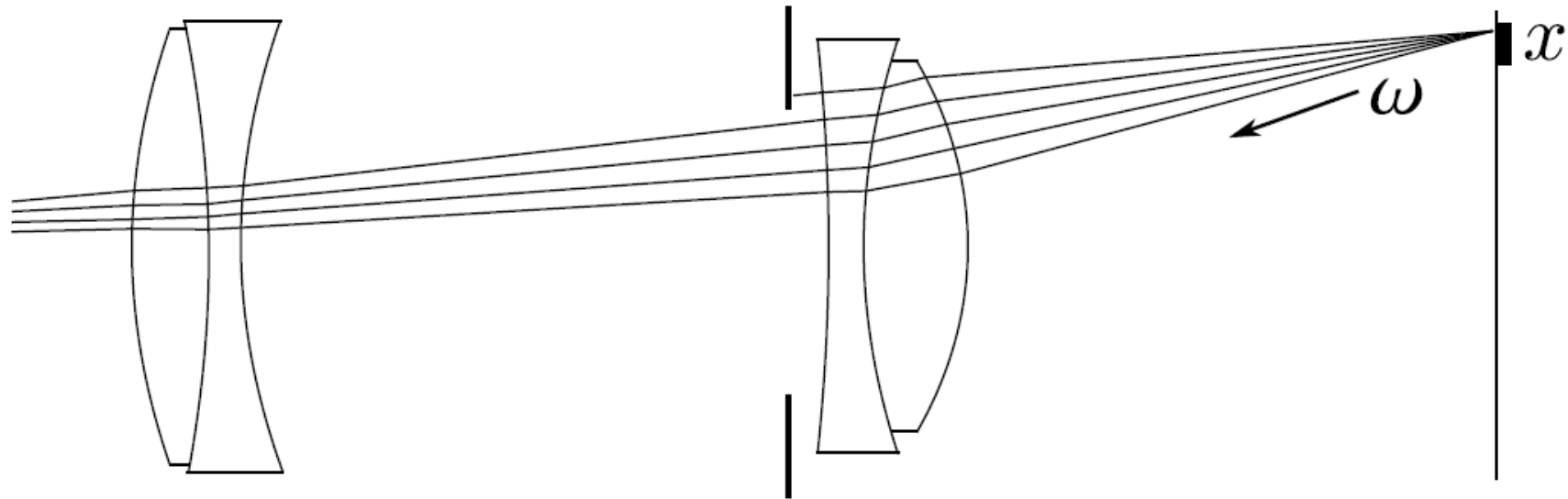
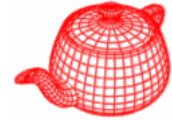
L : radiance

T : image to object space transformation

S : shutter function

P : sensor response characteristics

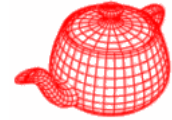
Measurement equation



$$R = \Delta t \cdot \int \int L(T(x, \omega)) \cos \theta \, dx \, d\omega$$

L : radiance T : image to object space transformation

Solving the integral



Problem: given a function f and domain Ω , how to calculate

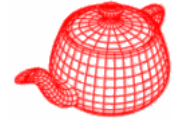
$$\int_{\Omega} f(x) dx$$

Solution: Monte Carlo method:

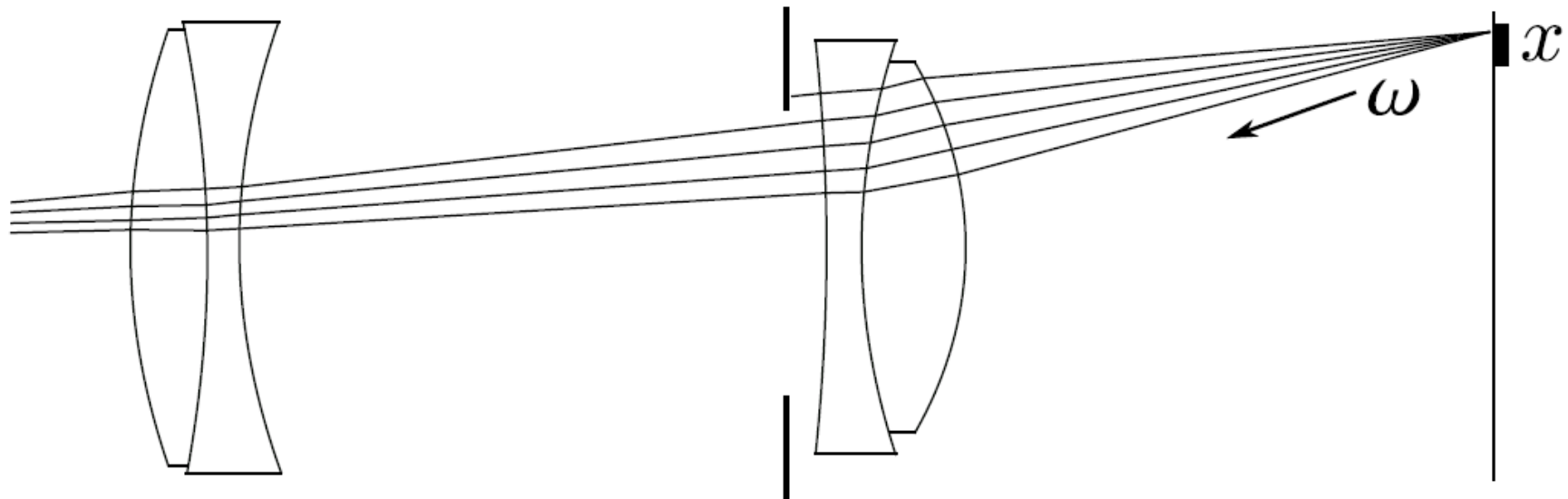
$$\int_{\Omega} f(x) dx \approx \left[\frac{1}{N} \sum_{i=1}^N f(x_i) \right] \cdot \int_{\Omega} dx$$

where x_1, x_2, \dots, x_N are uniform distributed random samples in Ω .

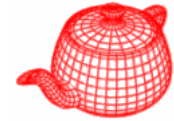
Algorithm



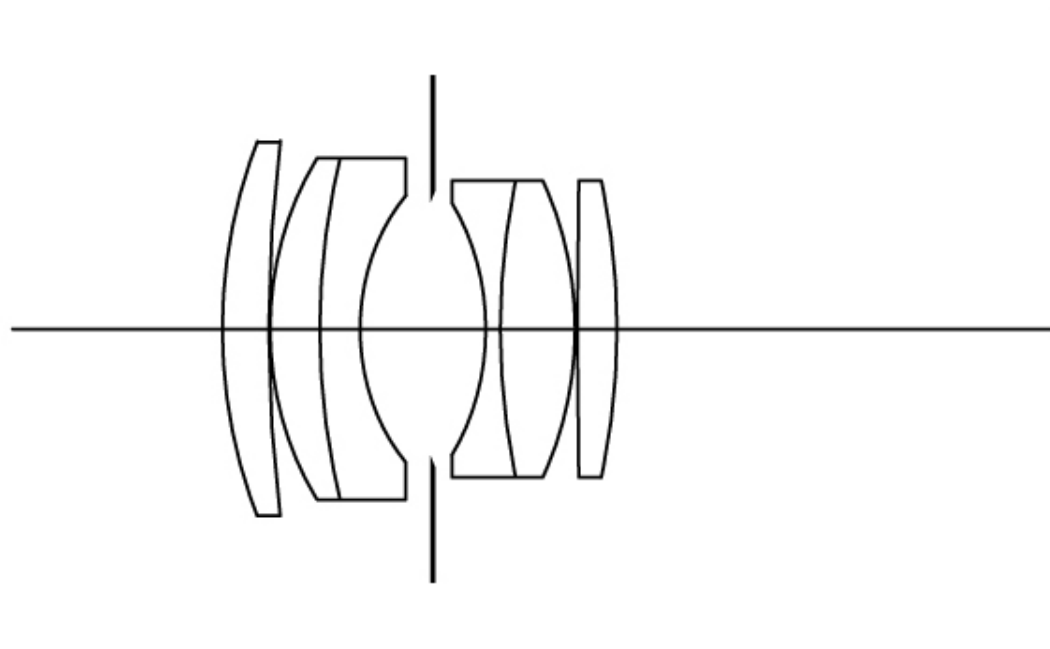
- 1 For each pixel on the image, generate some random samples x_i and ω_i uniformly.
- 2 For each x_i and ω_i , calculate $T(x_i, \omega_i)$.
- 3 Shoot the ray according to the result of $T(x_i, \omega_i)$ into the scene, and calculate the radiance.
- 4 Set the pixel value to the average of radiance.



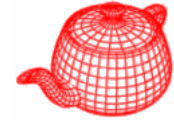
Tracing rays through lens system



- ① $R = \text{Ray}(x_i, \omega_i)$
- ② Calculate the intersection point p for each lens element E_i from rear to front.
 - ① Return zero if p is outside the aperture of E_i .
 - ② Compute the new direction by Snell's law if the medium is different.

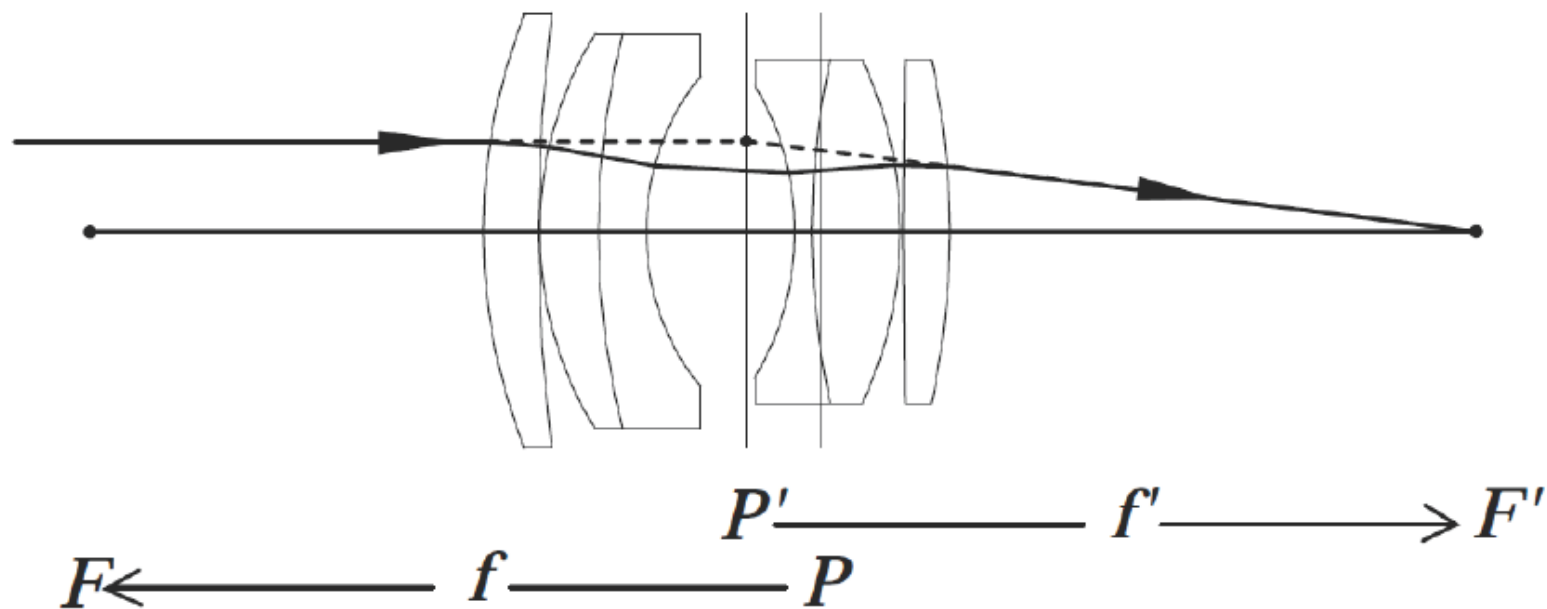
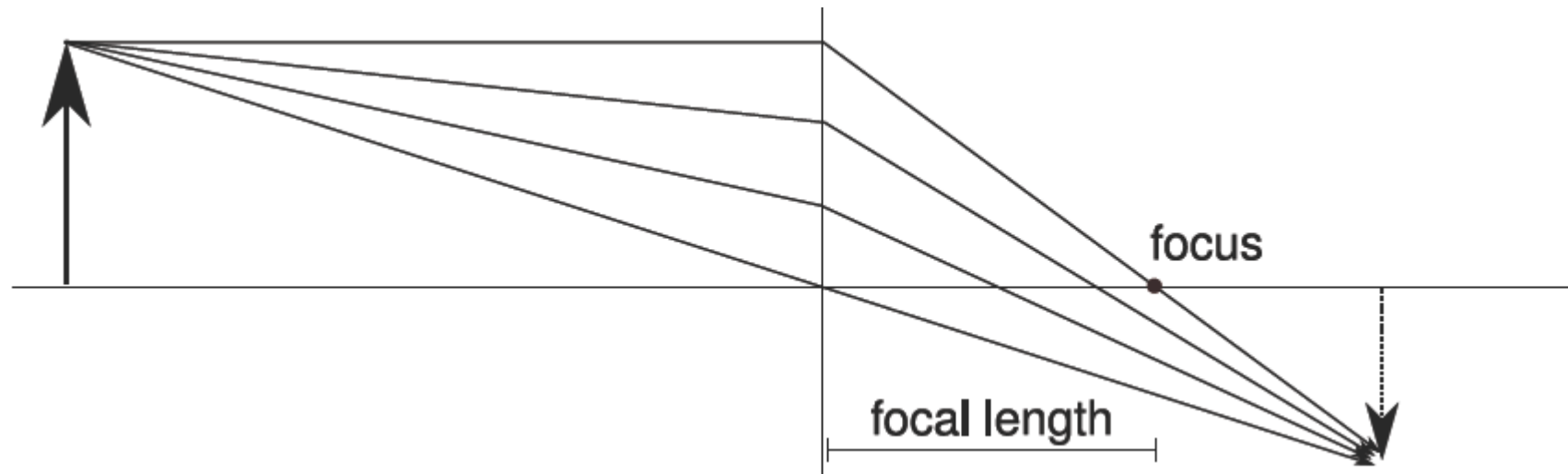
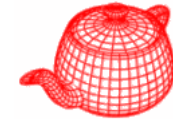


Ideal lens approximation

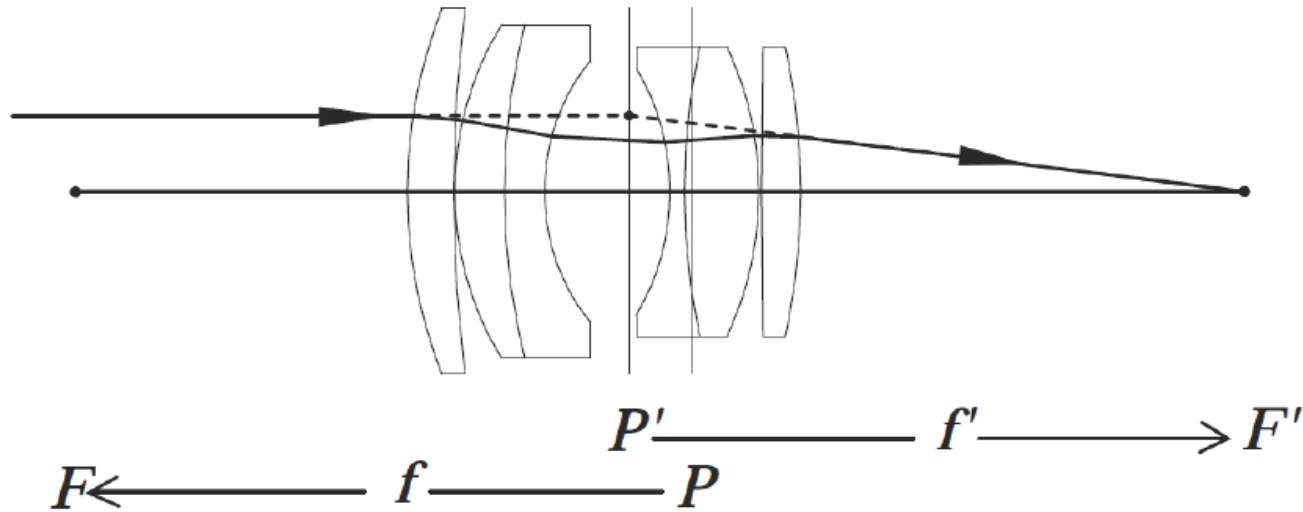
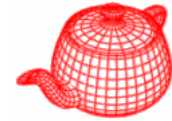


- In some situations we need an ideal lens approximation.
 - ▶ Ideal lens: each point in object space is imaged onto a single point in the image space.
 - ▶ All points on the **plane of focus** map onto the image plane.
- Thin lens approximation assumes that the thickness of lens is zero.
- Thick lens approximation has additional parameter of thickness.

Thin lens and thick lens



Finding thick lens approximation

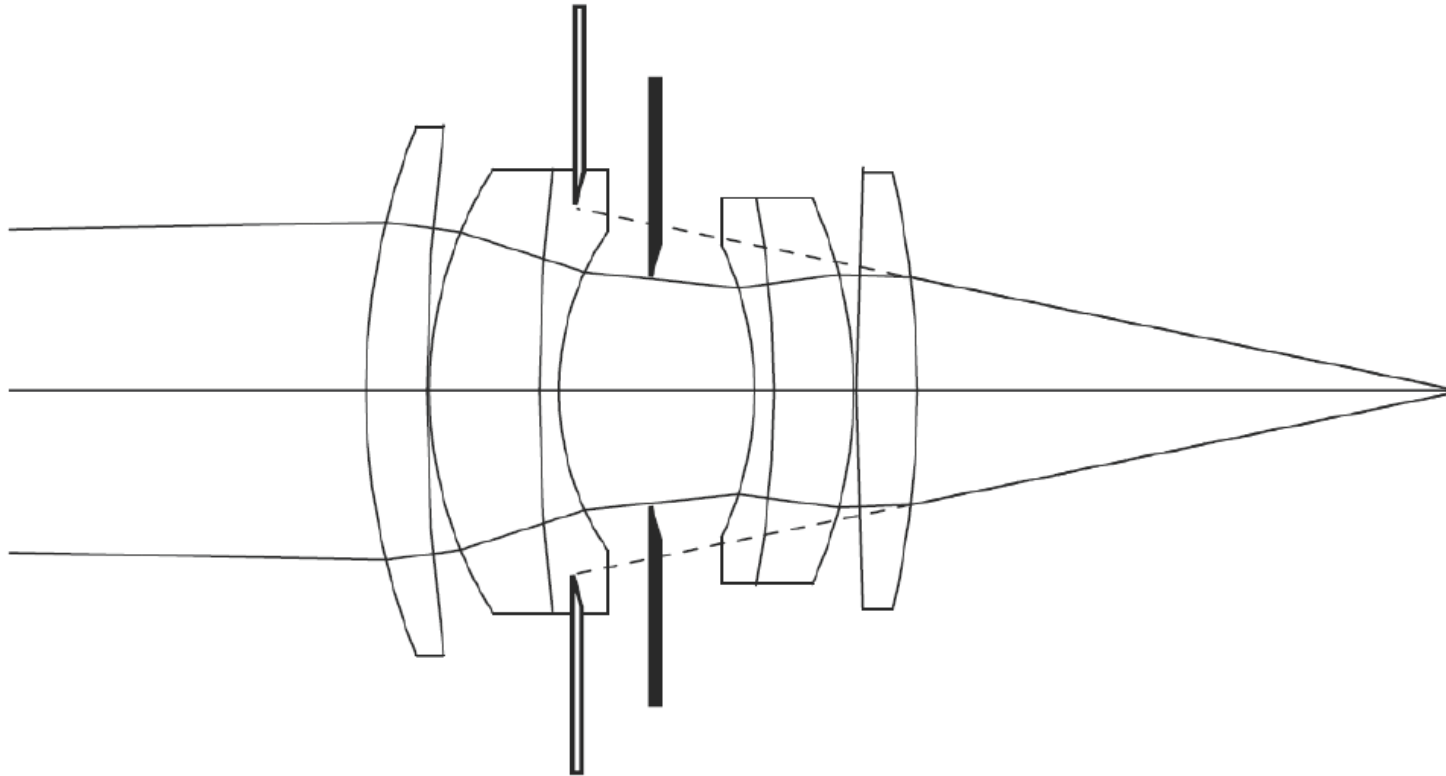
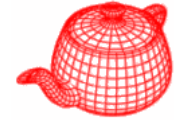


- 1 Shoot a ray parallel to the axis to find the focus.
- 2 Find the principal plane by intersecting the refracted ray and parallel one.
- 3 Find the secondary principal plane by tracing from another side.

Applications of thick lens approximation

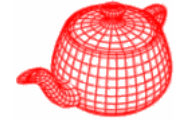
- Faster way to calculate the transform
- Autofocus
- Calculate the exit pupil

Exit pupil



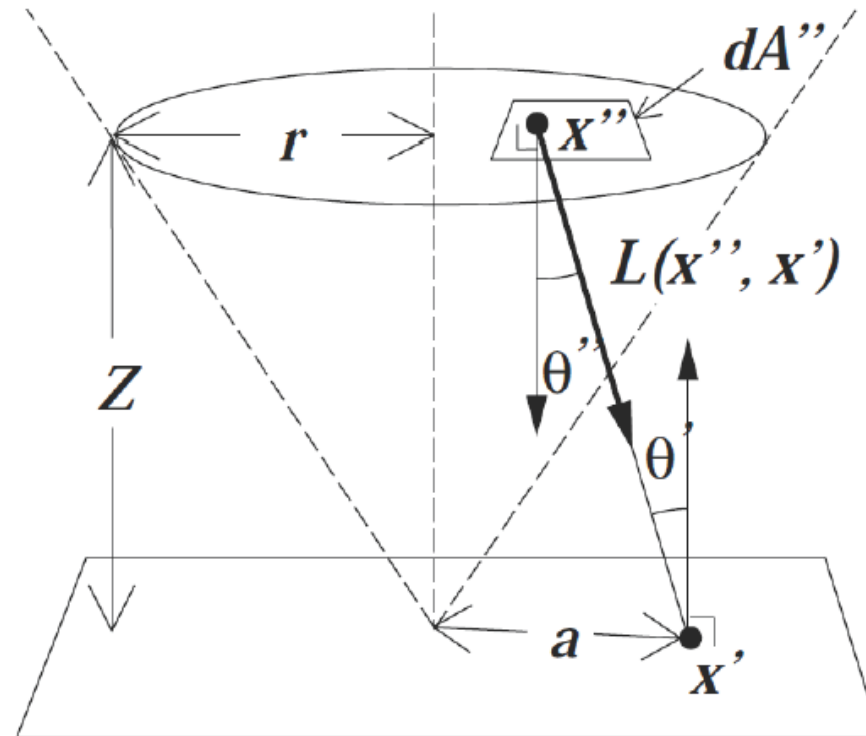
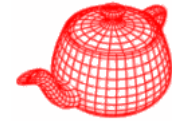
The exit pupil is the effective aperture stop in the image space which allows ray incidence.

Finding exit pupil



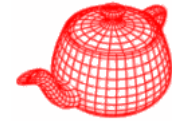
- Finding the exit pupil:
 - ① For each aperture stop, calculate its image by thick lens approximation.
 - ② Find the aperture stop whose image subtends the smallest solid angle.
- You may also use the aperture of the nearest lens as the exit pupil.

Integral over the exit pupil

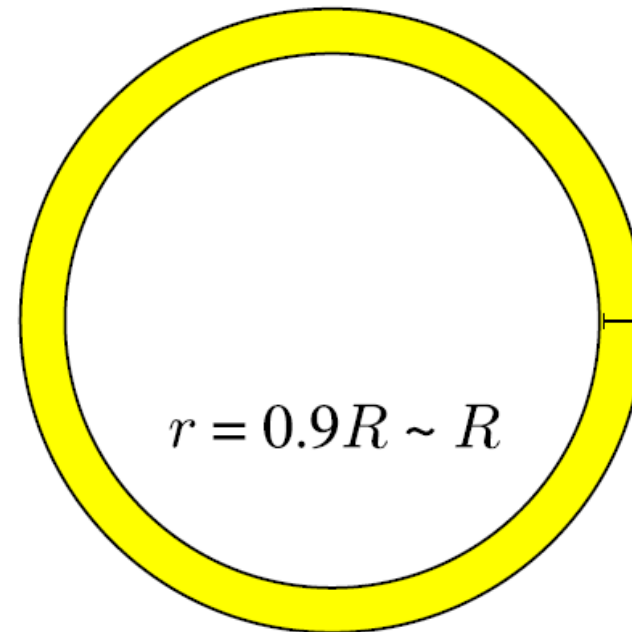
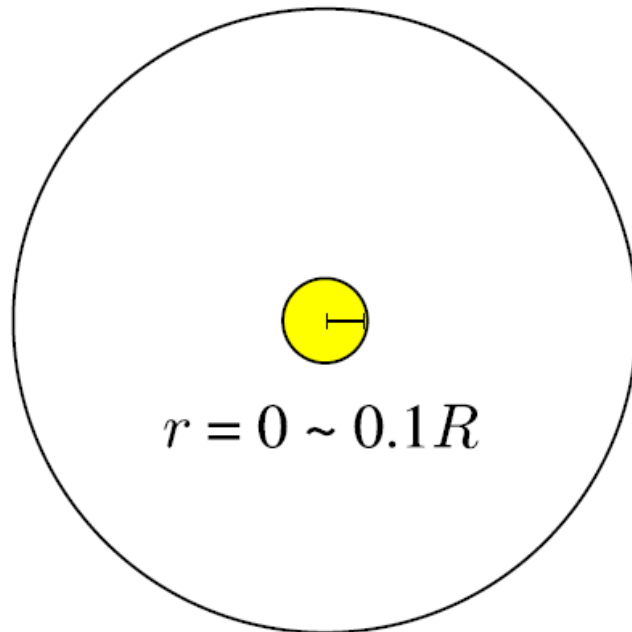


$$E(x') = \frac{1}{Z^2} \int_{x'' \in D} L(x'', x') \cos^4 \theta' dA''$$

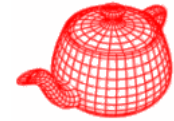
Sampling a disk uniformly



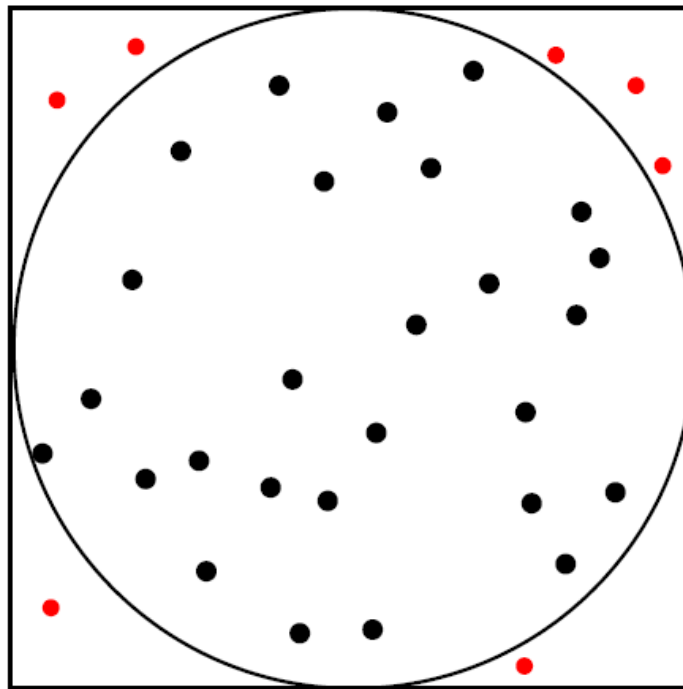
- Now we need to obtain random samples on a disk uniformly.
- How about uniformly sample r in $[0, R]$ and θ in $[0, 2\pi]$ and let $x = r \cos \theta, y = r \sin \theta$?
 - ▶ The result is not uniform due to coordinate transformation.



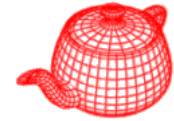
Rejection



- 1 Uniformly sample a point in the bounding square of the disk.
- 2 If the sample lies outside the disk, reject it and sample another one.



Another method



- Sample r and θ in a specific way so that the result is uniform after coordinate transformation.

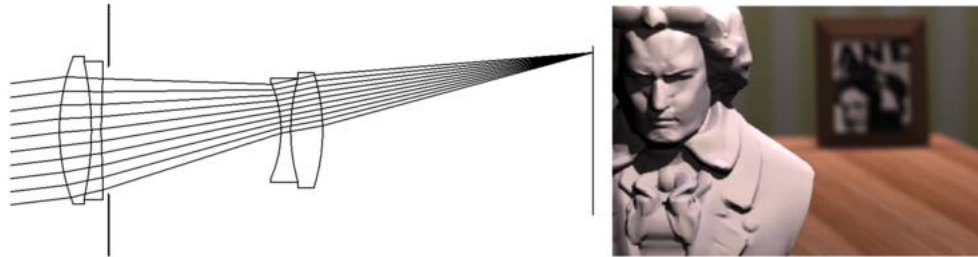
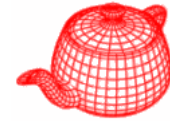
- Let

$$r = \sqrt{\xi_1}, \theta = 2\pi\xi_2$$

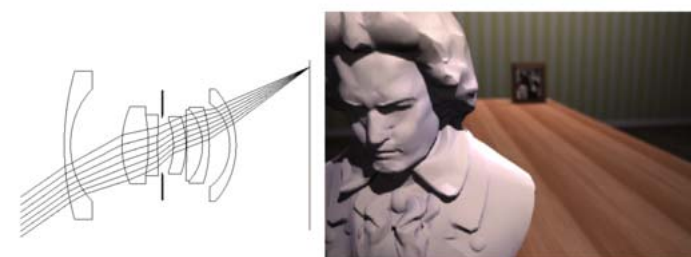
where ξ_1 and ξ_2 are random samples distributed in $[0, 1]$ uniformly.

- This produce uniform samples on a disk after coordinate transformation. We will prove it later in chapter 14 “Monte Carlo integration”.

Ray Tracing Through Lenses



200 mm telephoto



35 mm wide-angle



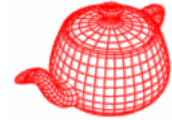
50 mm double-gauss



16 mm fisheye

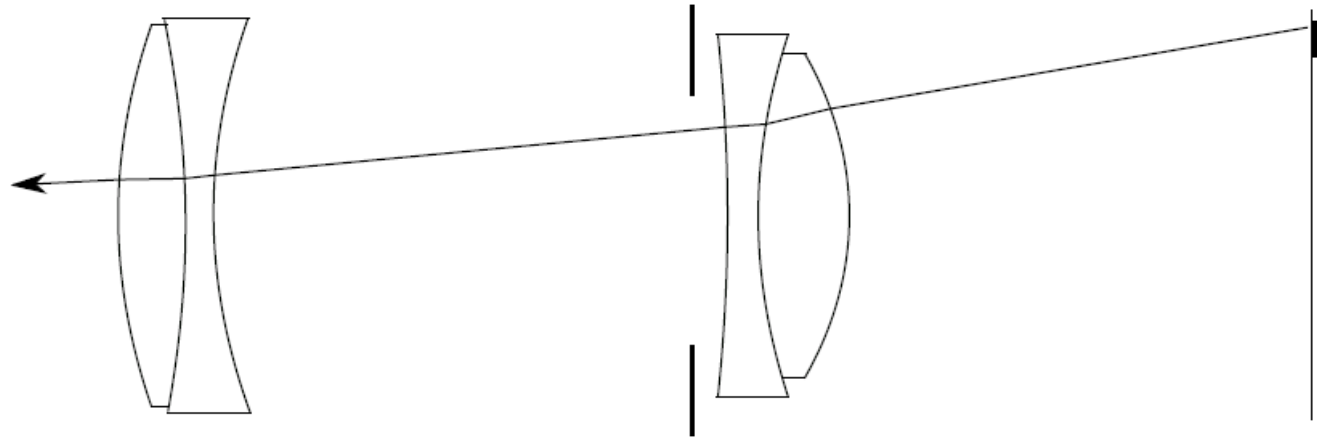
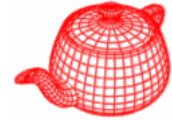
From Kolb, Mitchell and Hanrahan (1995)

Assignment #2



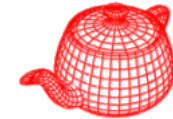
- Write the “realistic” camera plugin for PBRT which implements the realistic camera model.
- The description of lens system will be provided.
- `GenerateRay(const Sample &sample, Ray *ray)`
 - ▶ PBRT generate rays by calling `GenerateRay()`, which is a virtual function of `Camera`.
 - ▶ PBRT will give you pixel location in `sample`.
 - ▶ You need to fill the content of `ray` and return a value for its weight.

Assignment #2

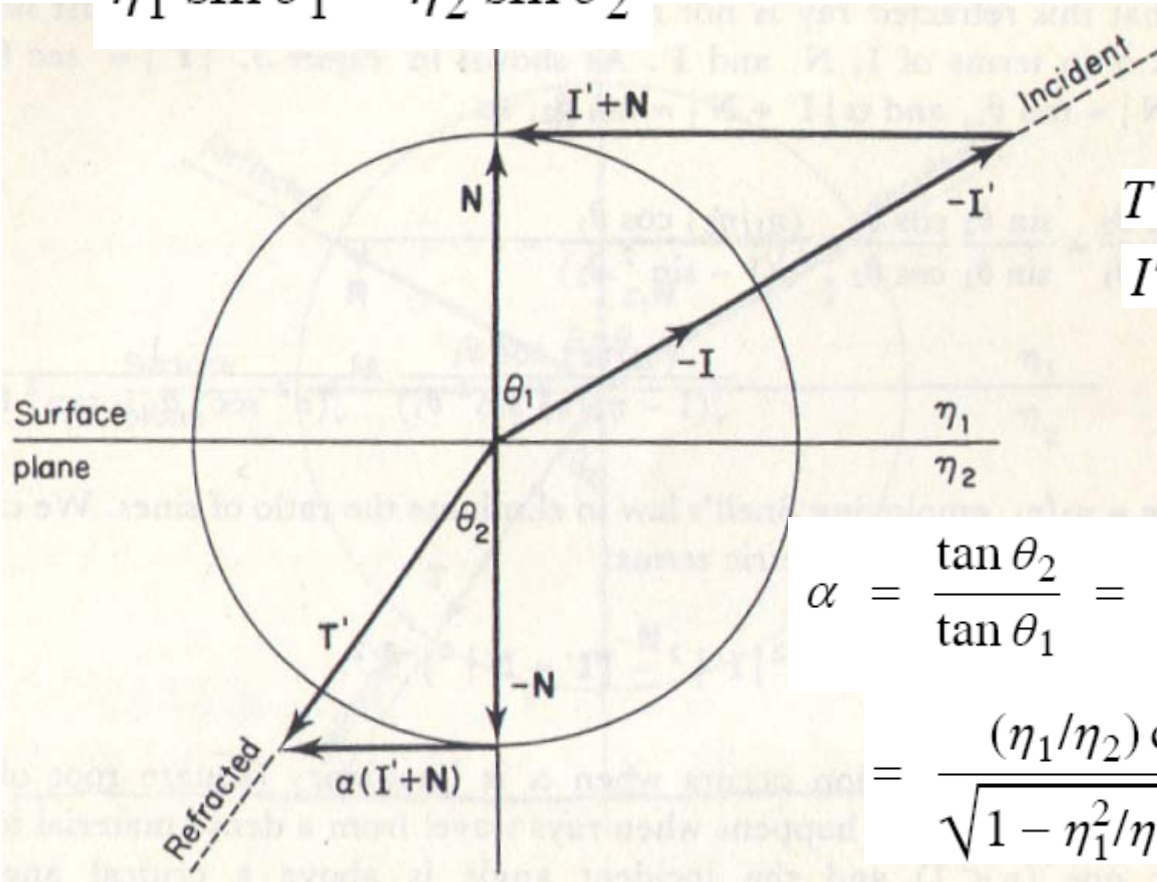


- 1 Sample a point on the exit pupil uniformly.
 - ▶ Hint: `sample.lensU` and `sample.lensV` are two random samples distributed in $[0, 1]$ uniformly.
- 2 Trace this ray through the lens system. You can return zero if this ray is blocked by an aperture stop.
- 3 Fill ray with the result and return $\frac{\cos^4 \theta'}{Z^2}$ as its weight.

Whitted's method



$$\eta_1 \sin \theta_1 = \eta_2 \sin \theta_2$$



$$T' = \alpha(I' + N) - N \text{ for some } \alpha$$

$$I' = I / (-I \cdot N)$$

$$|I' + N| = \tan \theta_1$$

$$\alpha |I' + N| = \tan \theta_2$$

$$\alpha = \frac{\tan \theta_2}{\tan \theta_1} = \frac{\sin \theta_2 \cos \theta_1}{\sin \theta_1 \cos \theta_2} = \frac{(\eta_1/\eta_2) \cos \theta_1}{\sqrt{1 - \sin^2 \theta_2}}$$

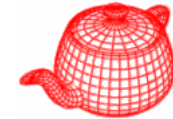
$$= \frac{(\eta_1/\eta_2) \cos \theta_1}{\sqrt{1 - \eta_1^2/\eta_2^2 \sin^2 \theta_1}} = \frac{1}{\sqrt{n^2 \sec^2 \theta_1 - \tan^2 \theta_1}}$$

$$\uparrow$$

$$|I'| = \sec \theta_1$$

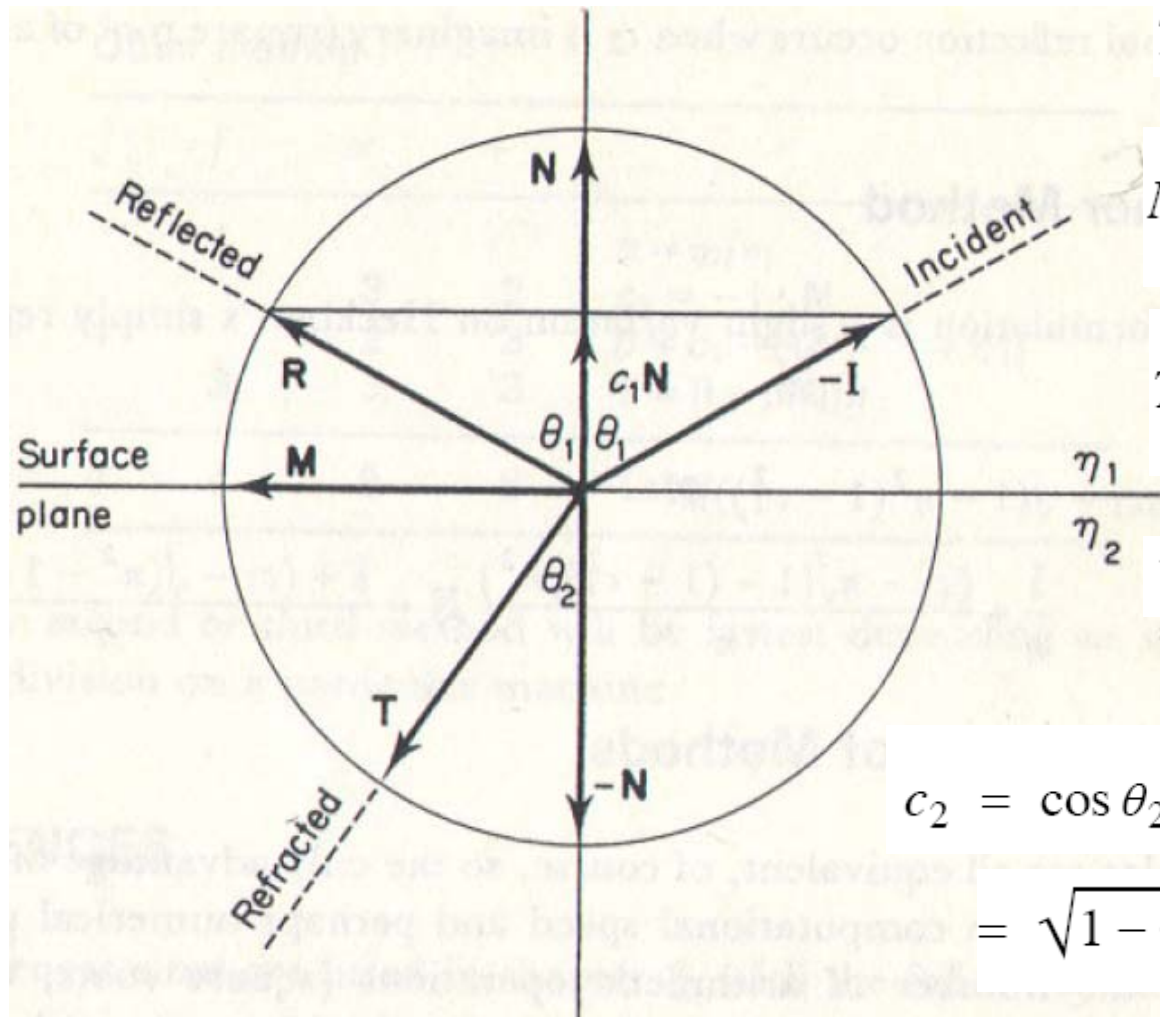
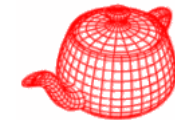
$$\alpha = (n^2 |I'|^2 - |I' + N|^2)^{-1/2}$$

Whitted's method



Whitted's Method				
$\sqrt{\quad}$	/	\times	+	
	1			$n = \eta_2 / \eta_1$
	3	3	2	$I' = I / (-I \cdot N)$
			3	$J = I' + N$
1	1	8	5	$\alpha = 1 / \sqrt{n^2(I' \cdot I') - (J \cdot J)}$
		3	3	$T' = \alpha J - N$
1	3	3	2	$T = T' / T' $
2	8	17	15	TOTAL

Heckber's method



$$T = \sin \theta_2 M - \cos \theta_2 N$$

$$M = \frac{I_{perp}}{|I_{perp}|} = \frac{I + c_1 N}{\sin \theta_1}$$

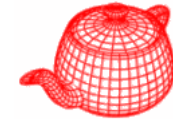
$$T = \frac{\sin \theta_2}{\sin \theta_1} (I + c_1 N) - \cos \theta_2 N$$

$$T = \eta I + (\eta c_1 - c_2) N$$

$$c_2 = \cos \theta_2 = \sqrt{1 - \sin^2 \theta_2}$$

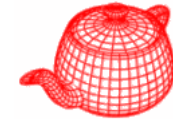
$$= \sqrt{1 - \eta^2 \sin^2 \theta_1} = \sqrt{1 - \eta^2 (1 - c_1^2)}$$

Heckbert's method



Heckbert's Method				
$\sqrt{\quad}$	/	\times	+	
	1			$\eta = \eta_1/\eta_2$
		3	2	$c_1 = -I \cdot N$
1		3	2	$c_2 = \sqrt{1 - \eta^2(1 - c_1^2)}$
		7	4	$T = \eta I + (\eta c_1 - c_2)N$
1	1	13	8	TOTAL

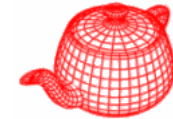
Other method



$$\begin{aligned}
 T &= \eta I + (\eta c_1 - \sqrt{1 - \eta^2(1 - c_1^2)})N \\
 &= \frac{I}{n} + \frac{c_1 - n\sqrt{1 - (1 - c_1^2)/n^2}}{n} N \\
 &= \frac{I + (c_1 - \sqrt{n^2 - 1 + c_1^2})N}{n}
 \end{aligned}$$

Other Method				
$\sqrt{\quad}$	/	\times	+	
	1			$n = \eta_2/\eta_1$
		3	2	$c_1 = -I \cdot N$
1		2	3	$\beta = c_1 - \sqrt{n^2 - 1 + c_1^2}$
	3	3	3	$T = (I + \beta N)/n$
1	4	8	8	TOTAL

Comparisons



Whitted's Method				
$\sqrt{\quad}$	/	\times	+	
	1			$n = \eta_2/\eta_1$
	3	3	2	$I' = I/(-I \cdot N)$
			3	$J = I' + N$
1	1	8	5	$\alpha = 1/\sqrt{n^2(I' \cdot I') - (J \cdot J)}$
		3	3	$T' = \alpha J - N$
1	3	3	2	$T = T'/ T' $
2	8	17	15	TOTAL

Heckbert's Method				
$\sqrt{\quad}$	/	\times	+	
	1			$\eta = \eta_1/\eta_2$
		3	2	$c_1 = -I \cdot N$
1		3	2	$c_2 = \sqrt{1 - \eta^2(1 - c_1^2)}$
		7	4	$T = \eta I + (\eta c_1 - c_2)N$
1	1	13	8	TOTAL

Other Method				
$\sqrt{\quad}$	/	\times	+	
	1			$n = \eta_2/\eta_1$
		3	2	$c_1 = -I \cdot N$
1		2	3	$\beta = c_1 - \sqrt{n^2 - 1 + c_1^2}$
	3	3	3	$T = (I + \beta N)/n$
1	4	8	8	TOTAL